Light – Matter Interactions (II)

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- Phenomenology of magnetic spectroscopies
- Electronic structure theory, linear-response theory
- Theory/understanding of magnetic spectroscopies
 - Optical regime
 - Ultraviolet and soft X-ray regime



UPPSALA In the beginning ... First observation of magneto-optics

Magneto-optical Faraday effect (1845)

Observation of interaction light-magnetism enormous impact on development of science!





Michael Faraday (1791 – 1868)





Magneto-optical Voigt effect

Voigt effect (1899)

Very different from Faraday effect; Voigt effect is "quadratic" (<u>even</u>) in M

Kerr effect



Imaging of magnetic domains using Voigt and Kerr effect in reflection



Woldemar Voigt (1850 – 1919)

Voigt effect



(Courtesy R. Schäfer)



Magnetic circular and linear dichroism





Development of light-magnetic material interaction





Observation of the spin Hall effect

Kato, Myers, Gossard & Awschalom, Science **306**, 1910 (2004)

Spin Hall effect



Dyakonov & Perel, JETP Lett. **13**, 467 (1971)

Material: non-magnetic n-GaAs [110] MO Kerr rotation detection $\sim 10^{-5}$ deg.





Spin Hall effect in heavy metals

Gives rise to spin-orbit torque

Direct observation of SHE in pure heavy-metal difficult because of short spin lifetime and spin diffusion length



MOKE detection could be possible due to penetration depth



Miron et al, Nature **476**, 189 (2011) Liu et al, Science **336**, 555 (2012)





Experimental direct observation of spin Hall effect Pt





Stamm, Murer, Berritta, Feng, Gabureac, Oppeneer & Gambardella, PRL **119**, 087203 (2017)

Excellent agreement with experiment Estimated $l_s=11.4\pm2$ nm for pure Pt

Accurate MOKE measurements of SH conductivity in heavy metals feasible with nrad sensitivity



Optically induced magnetization



Kimel et al, Nature **435**, 655 (2005)

Due to nonlinear "opto-magnetic" effect, the inverse Faraday effect:

Induces magnetization δM

$$\vec{M}^{ind} \propto v_{IFE} \cdot E_i E_j^*$$

Could potentially lead to a fast, optically driven magnetization reversal



All-optical writing of magnetic domains

After exposure



Stanciu et al, PRL 99, 047601 (2007)

 Due to *inverse Faraday effect?* Background all-optical magn. recording
 Erasing & writing with fs-laser pulses
 Approx. 10³ times faster recording? (symposium Th. Rasing, A. Kirilyuk)



Lambert et al, Science 345, 1337 (2014)





Ultrafast magnetism

Measurement of ultrafast magnetic response with time-resolved magneto-optics



Use 2nd level: Combination of Maxwell-Fresnel theory and *ab initio* quantum theory

Theoretical description of light – magnetism interaction

Geometry & materials' boundary conditions:

Fresnel equation for modes in material:

$$\left\lfloor n^2 \mathbf{1} - \boldsymbol{\epsilon} - \boldsymbol{n} : \boldsymbol{n} \right\rfloor \cdot \boldsymbol{E} = 0$$

$$\begin{pmatrix} E_{s}^{r} \\ E_{p}^{r} \end{pmatrix} = \begin{pmatrix} r_{ss} & r_{sp} \\ r_{ps} & r_{pp} \end{pmatrix} \bullet \begin{pmatrix} E_{s}^{i} \\ E_{p}^{i} \end{pmatrix}$$

$$E_{p}$$
 e b e

$$\frac{E_s^r}{E_s^i} \equiv r_{ss} = \frac{n_i \cos \theta_i - n_t \cos \theta_t}{n_i \cos \theta_i + n_t \cos \theta_t}$$
$$\frac{E_s^t}{E_s^i} \equiv t_{ss} = \frac{2n_i \cos \theta_i}{n_i \cos \theta_i + n_t \cos \theta_t}$$

Polarization analysis or intensity measurement

And: *ab initio* theory for calculation of $\varepsilon(\omega)$



Dependence of the dielectric tensor on fields

The dielectric tensor depends on external fields $\vec{\varepsilon} \rightarrow \vec{\varepsilon}(\vec{k}, \omega, \vec{B}, \vec{E})$

Use a Taylor expansion for effects to lowest order:



All the (linear) phenomena can be described, using the Fresnel formalism

 $(\vec{\mu} \rightarrow 1)$



Magnetic effects in Fresnel equations





Magn. effects probe always $\sim M$ or $\sim M^2$ (to lowest order)

Examples:

$$MCD = \frac{\omega d}{2c} \operatorname{Re}\left[\frac{\varepsilon_{xy}}{n}\right] \text{ Odd in M; } MLD = \frac{\omega d}{2cn} \operatorname{Im}\left[\varepsilon_{\parallel} - \varepsilon_{\perp} - \frac{\varepsilon_{xy}^{2}}{\varepsilon_{\perp}}\right] \text{ even}$$
$$\varepsilon_{\alpha\alpha} \approx \varepsilon_{\alpha\alpha}^{0} + \zeta_{\alpha} M_{\alpha}^{2}$$



Magneto-optical Kerr and Faraday effects



polar Kerr effect, normal incidence

Faraday effect, normal incidence



Linear (odd) in M spectroscopies:



Classification criteria:

- 1. Magnetic parity
- 2. Transmission or reflection
- 3. Polarization or intensity
- 4. Linearly or circ. polarized light
- Suitable for (element-selective) study of ferro-, ferrimagnets



Quadratic (even) in M spectroscopies:





 Suitable for (element-selective) study of antiferromagnets (and ferromagnets as well)



Lifetime broadening happens and needs to be taken into account



Lifetime broadening – linear magneto-optics



Optical frequencies: lifetime $\Gamma=1/\tau \approx 0.03$ Ry

Oppeneer, Handbook of Magnetic Materials, Vol. 13 (2001)





Effective Kohn-Sham Hamiltonian:

$$\hat{H} = \begin{bmatrix} -\frac{\nabla^2}{2m} + V_{e,N}(\vec{r}) + V_0(\vec{r}) \end{bmatrix} 1 + \vec{B}_{xc}(\vec{r}) \cdot \hat{\vec{\sigma}} + \xi \hat{\vec{\ell}} \cdot \hat{\vec{\sigma}}$$
Spin-density (2x2):
$$n(\vec{r}) = \{n_0(\vec{r})1 + \vec{m}(\vec{r}) \cdot \vec{\sigma}\}/2$$

$$n(\vec{r}) = u_0\{n, (\vec{r}) - n_1(\vec{r})\}$$

 $m(\vec{r}) = \mu_B \left\{ n_{\uparrow}(\vec{r}) - n_{\downarrow}(\vec{r}) \right\}$

Vary the two magnetic interactions (exchange & spin-orbit) to deduce how magnetic spectra depend on these.



Effect of SOI and exchange interaction





Leading quantity determining the valence band MO effect is spin-orbit coupling ξ (**L.S**) \Rightarrow Kerr and Faraday effect scale linear in the SOC, not in the exc.-splitting!



UPPSALA What about the X-ray regime ?

X-ray magnetic circular dichroism $\vec{E}_{\pm}(z,t) \propto (\vec{e}_x \pm i\vec{e}_y) \cdot e^{i\omega/c(n_{\pm}\cdot z) - i\omega t}$



Understand origin of and perform *ab initio* calculations for XMCD & XMLD at L-edges



Note on importance of XMCD – sum rules



Thole et al, PRL **68**, 1943 (1992) Carra et al, PRL **70**, 694 (1993) **Sum rules** $/\mu_{B} \sim \frac{A_{L3} - 2A_{L2}}{A_{Iso}} \qquad \text{Atomic spin moment}$ $/\mu_{B} \sim \frac{A_{L3} + A_{L2}}{A_{Iso}} \qquad \text{Atomic orbital moment}$

The XMCD sum rules are not exact but are intensively used, because they allow an element-selective determination of the spin & orbital moment on a 3d element in a material.

(Lecture E. Goering)







Basic electronic structure picture

1) Spin-splitting of 3d states due to exchange interaction

2) Helicity-dependent optical selection rules

$$left: \Delta m_{\ell} = -1, \quad \Delta m_{s} = 0$$
$$right: \Delta m_{\ell} = +1, \quad \Delta m_{s} = 0$$



Leads to different absorption of left/right circ. pol. radiation (trans. probabilities)



Ab initio calculated XMCD spectra - Effect of $\Delta_{\!xc}$

Very different size of SOI and $\Delta_{\rm xc}$!

XMCD
$$\approx$$
 Im $[\mathcal{E}_+ - \mathcal{E}_-]$

To lowest order the XMCD does not depend on Δ_{ex} :



Many calculations ignore Δ , but there is a small effect !



Kunes et al, PRB 64, 174417 (2001)



Comparison with experimental XMCD spectra



 $n_{\pm} = 1 - (\delta \pm \Delta \delta) + i(\beta \pm \Delta \beta)$

Exchange-split core states give somewhat better results when compared to experimental spectra !

Kunes et al, PRB 64, 174417 (2001)





Magnetovolume effect? Why much smaller than MCD? Spin-orbit interaction? $n_{\parallel} \neq n_{\perp}$ Cf. Faraday, Kerr: linear in ξ_{so}

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 $MLD \propto (I_{\parallel} - I_{\perp})/(I_{\parallel} + I_{\perp})$



Measurements & ab initio calculations XMLD





Further results of *ab initio* calculations XMLD



> small effect ~ 5% $A_{XMLD} \approx \frac{\omega d}{c} \operatorname{Im}[n_{\parallel} - n_{\perp}] \approx \frac{\omega d}{2cn} \operatorname{Im}[\varepsilon_{\parallel} - \varepsilon_{\perp}]$ > good *ab initio* theory

Why do we have these spectral structures?



Simple model for X-ray magnetic spectroscopies

Core-states are (k) dispersionless

$$\boldsymbol{\epsilon}_{xy}^{(1)}(\boldsymbol{\omega}) = \frac{4\pi^2 e^2}{\hbar V_{uc}} \operatorname{Im} \sum_{\boldsymbol{k}} \sum_{\substack{c \text{ occ.}\\n \text{ un.}}} r_{nc}^x(\boldsymbol{k}) r_{cn}^y(\boldsymbol{k}) \,\delta(\boldsymbol{\omega} - \boldsymbol{\omega}_{nc}(\boldsymbol{k}))$$

Expand ϵ_{xy} considering 2p -> 3d transitions

Selection rules on m :

$$\begin{split} \varepsilon_{+} &= \varepsilon_{xx} + i\varepsilon_{xy} \Longrightarrow \Delta m = +1 \\ \varepsilon_{-} &= \varepsilon_{xx} - i\varepsilon_{xy} \Longrightarrow \Delta m = -1 \\ \varepsilon_{0} &= \varepsilon_{\parallel} \Longrightarrow \Delta m = 0 \end{split}$$

X-ray spectroscopies:
XAS
$$\approx \text{Im}[2\varepsilon_{\perp} + \varepsilon_{\parallel}]/3$$

Sum of all transitions Δm
XMCD $\approx \text{Im}[\varepsilon_{+} - \varepsilon_{-}]$
Difference of transitions
with Δm =+1 and Δm =-1
XMLD
 $\approx \text{Im}[\varepsilon_{\parallel} - \varepsilon_{\perp}] = \text{Im}[\varepsilon_{\parallel} - (\varepsilon_{+} + \varepsilon_{-})/2]$
Difference of transitions with
 Δm =0 and aver. Δm =+1 & -1

Understanding the shape of XMLD and XMCD

Develop model theory and perform ab initio calculations

Model theory (2p core):

- → Neglect SO in valence states (~ meV)
- → Consider only 2p ->3d transitions
- \rightarrow Expand ε functions with respect to Δ_{ex}

$$\operatorname{Im}[\varepsilon_{\mu}(\omega)] \propto \sum_{m,s} a^{\mu}_{\gamma,s}(j) D_{ms}\left(\omega \pm \frac{\gamma \Delta}{2}\right), \quad \gamma/2 = m - \mu + s$$
$$D_{ms} \text{ m- and spin-dependent 3d partial DOS}$$

 $\Delta_{\rm SO} (15 \text{ eV}) >> \Delta_{\rm ex} (1-3 \text{ eV}) >> \Delta_{\rm ex} (0.1-0.3 \text{ eV}) > \xi_{\rm so-v} (0.09 \text{ eV})$

XMLD
$$\approx \operatorname{Im}[\varepsilon_{\parallel} - \varepsilon_{\perp}] = \operatorname{Im}[\varepsilon_{\parallel} - (\varepsilon_{+} + \varepsilon_{-})/2]$$

Difference of transitions with $\Delta m=0$ and aver. $\Delta m=+1$ & -1



Model theory, results

 $D_{\rm ms}$: m and s-dependent partial DOS of unoccupied 3d states

$$\mathbf{XAS} \begin{cases} \approx 4 \sum_{m} (D_{m\uparrow} + D_{m\downarrow}) & j = 3/2 \\ \approx 2 \sum_{m} (D_{m\uparrow} + D_{m\downarrow}) & j = 1/2 \end{cases}$$
$$\mathbf{XMCD} \begin{cases} \approx +2 \sum_{m} (D_{m\uparrow} - D_{m\downarrow}) & j = 3/2 \end{cases}$$

m-orbital degeneracy:

XMLD $\approx \Delta \frac{d}{dE} (D_{\uparrow} - D_{\downarrow}) \qquad j = 1/2, \ 3/2$ $\approx \mp \Delta \frac{d}{dE} \{XMCD\}$ XAS branching ratio: 2/3 no magnetism (M invariant)

 $\begin{array}{c} \longrightarrow \\ L_2, L_3 \text{ equal & opposite} \\ \text{odd in } M, \text{ no f.o. effect of } \Delta \end{array}$

Small signal, proportional Δ !
related to energy deriv. XMCD
even in M (Δ is odd, D is odd)



Experimental check of XMCD-XMLD relation

m-orbital degeneracy: No crystal field, amorphous FeCo alloy

Spin-pol. unocc. 3d DOS $\left(D_{\uparrow} - D_{\downarrow} \right)$

Leading quantity Δ_{ex} is very small, yet crucial!



relation between XMCD-XMLD verified

Kunes et al, JMMM 272, 2146 (2004)



Explanation of the XMCD shape



Leading quantity for XMCD: exchange splitting of 3d DOS, determines XMCD shape



Explanation of the XMLD shape





How good are the assumptions in the model ?

XMCD and XMLD !
 Useful for studying the origin of XMCD and XMLD

 \rightarrow It is excellent for XAS,



Kunes & Oppeneer, PRB 67, 024431 (2003)

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Magnetocrystalline anisotropy in XMLD



 different combination of m-partial DOS probed, depending on M axis
 large magnetocrystalline anisotropy appears in XMLD spectra



Kunes & Oppeneer, PRB 67, 024431 (2003)



How about the 3p (M) semi-core edges?





Measured as intensity change of lin. polarized light in reflection (cf. Fresnel theory)

$$A = 2 \operatorname{Re}\left[\frac{\sin 2\theta_i \,\epsilon_{xy}}{n^4 \cos \theta_i^2 - n^2 + \sin \theta_i^2}\right] = 2 \operatorname{Re}\left[F(\theta, n) \,\epsilon_{xy}\right] = 2 \operatorname{Re}\left[F(\theta, n)\right] \operatorname{Re}\left[\epsilon_{xy}\right] - 2 \operatorname{Im}\left[F(\theta, n)\right] \operatorname{Im}\left[\epsilon_{xy}\right]$$

Demagnization mechanism: Compute *ab initio* ε_{xy} for several cases: 1) frozen magnon excitations, 2) reduced *exchange splitting* (spin-flips), 3) increased electron temperature T_e - construct the <u>change</u> in A(t) wrt $A(t=0) \rightarrow$ least square fit with experiment



$$A(t) = \frac{R(M+) - R(M-)}{R(M+) + R(M-)}$$

Turgut et al, PRB **94**, 220408R (2016)



Comparison of experiment and *ab initio* theory



For Co > Surprisingly small contribution from spin-flips (exch. split reduction) > Larger effect (2/3) is due to fast magnon excitation => reduction of M_z



Quadratic in *M* effect in-near-normal reflection



Schäfer-Hubert effect (or Voigt effect in reflection)





Near-normal incidence detection at Fe 3p edges

$$\theta_{\rm SH} \approx {\rm Re} \bigg[\frac{(n_{\parallel} - n_{\perp})n_0}{n_{\parallel}n_{\perp} - n_0^2} \bigg] \approx {\rm Re} \bigg[\frac{(\boldsymbol{\epsilon}_{\parallel} - \boldsymbol{\epsilon}_{\perp})n_0}{(n^2 - n_0^2)n} \bigg]$$



Comparison to *ab initio* calculations



Valencia et al, PRL **104**, 187401 (2010)

Importance of treating exchange splitting and spinorbit interaction on equal footing Importance of *hybridization* of j_z states

2p exchange splitting



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*j*_z-hybridization 3p semi-core level of Fe

Strong mixing of j, j_z states, SO splitting & exchange splitting equally large



No expansion in small quantity possible !



Magnetic spectroscopy is a highly sensitive tool that can detect minute magnetizations (spin Hall effect)

Exchange and spin-orbit splitting work together in different ways in valence and X-ray regime to bring about light - magnetic matter interaction

Ab initio quantum theory (effective single particle theory) works well but it is needed to know about its limitations

Current frontlines:

- 1) Ultrasensitive measurements to observe very small spin-orbit related effects (e.g. Inverse spin galvanic effect)
- 2) Ultrafast limit of modifications & control of magnetization, experiments and suitable theory
- 3) Nonlinear magneto-optic effects



- S. W. Lovesey and S. P. Collins, *X-Ray Scattering and Absorption by Magnetic Materials* (Clarendon Press, Oxford, 1996).
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- J. Stöhr and H.-C. Siegmann, *Magnetism: From Fundamentals to Nanoscale Dynamics* (Springer, Berlin, 2007).





Other way of describing effects – scattering formalism:





Deficiency of DFT-LDA for localized 4f states





Nearly localized f state => LDA+U better



Practicals' problem:

- 1) Material with magnetization in the scattering plane
- 2) Lin. pol. light *E*-vector at 45° to the magnetization
- 3) Consider *R(+M)-R(-M)*



Use the reflection coefficients to show that *R(+M)-R(-M)* is a measure of the magnetization and derive an expression for the magn. asymmetry:

$$A = \frac{R(+M) - R(-M)}{R(M) + R(-M)}$$