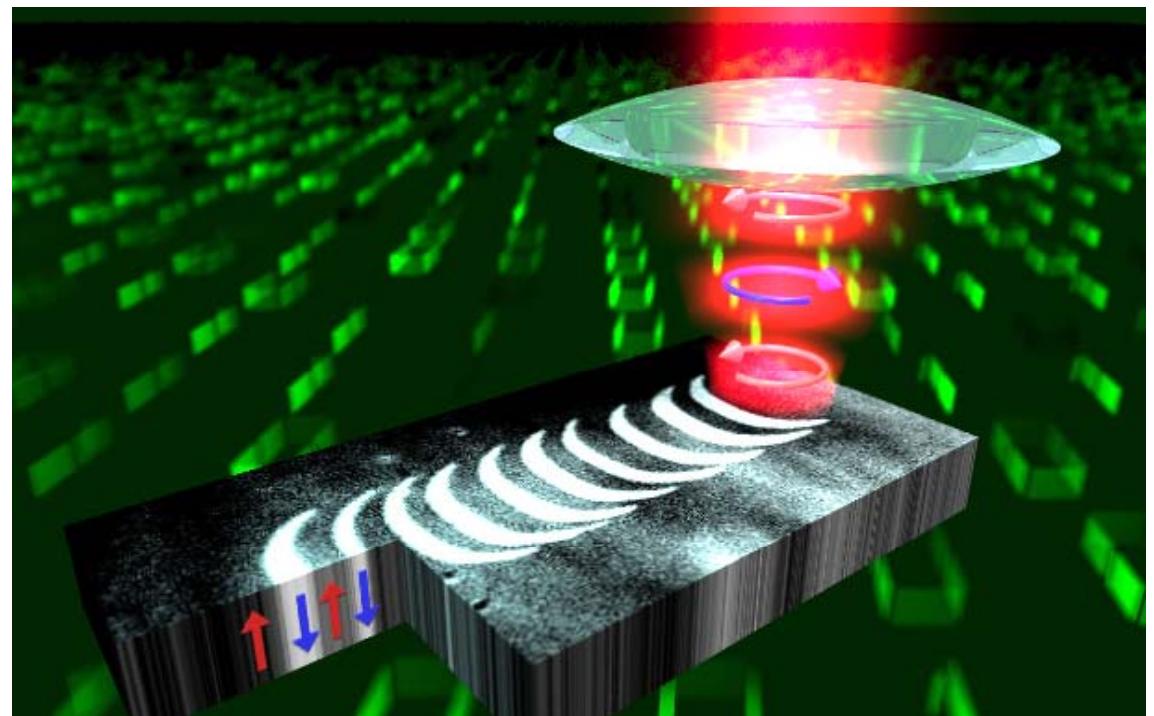
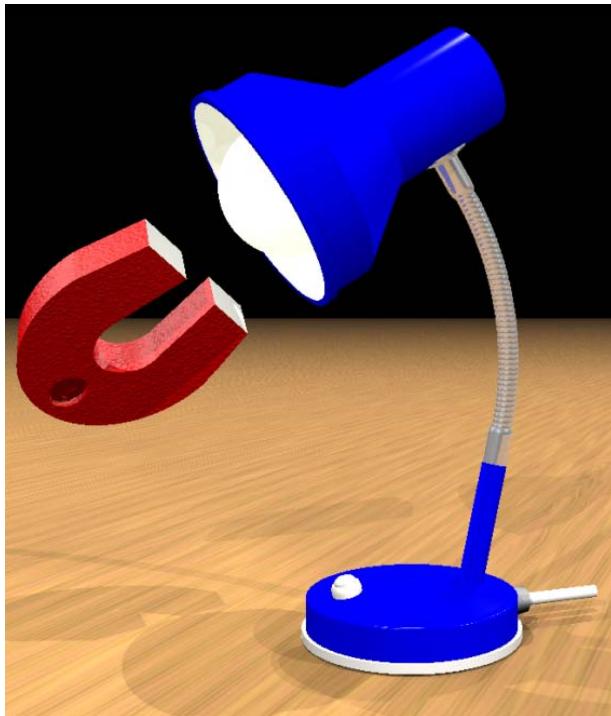


From magneto-optics to ultrafast manipulation of magnetism

Andrei Kirilyuk

Radboud University Nijmegen, The Netherlands



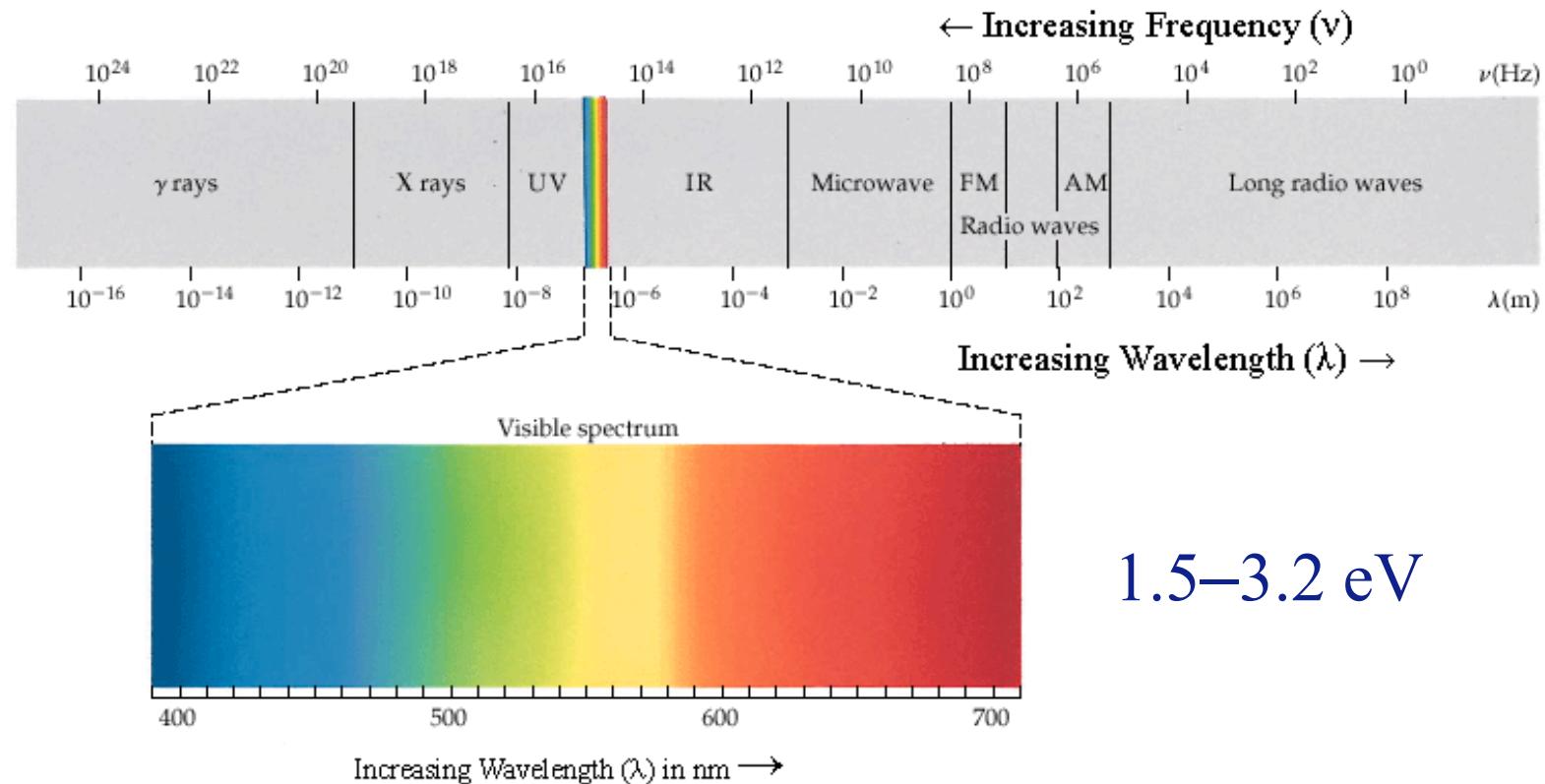
Outline of the lecture

- **Light as a probe**
 - linear magneto-optics
 - nonlinear (magneto-)optics
- **Example: all-optical FMR**
- **Light as an excitation**
 - classification of effects
 - basics of opto-magnetism
 - coherent control
 - local control of spins
- ***can this become too-ultrafast?***

Outline of the lecture

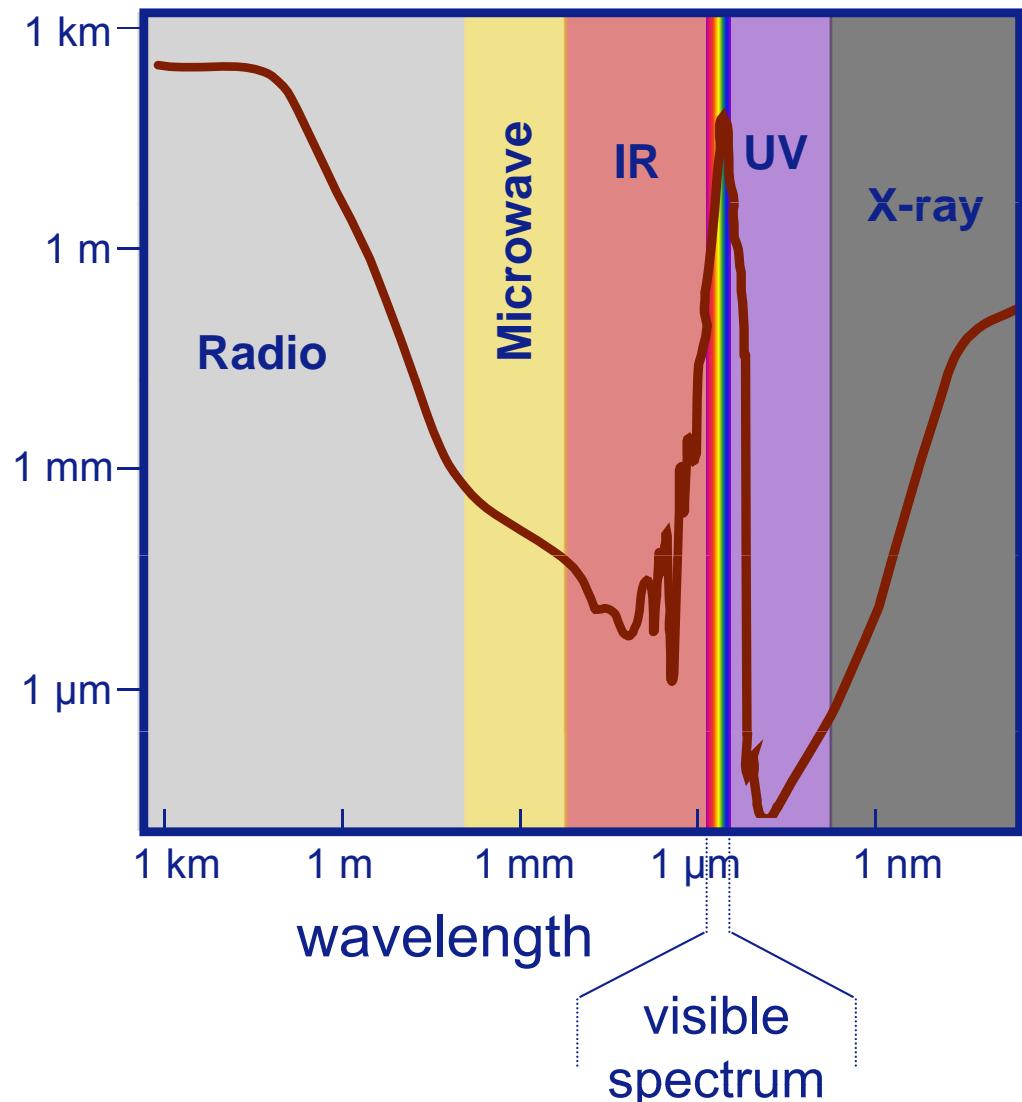
- **Light as a probe**
 - linear magneto-optics
 - nonlinear (magneto-)optics
- Example: all-optical FMR
- **Light as an excitation**
 - classification of effects
 - basics of opto-magnetism
 - coherent control
 - local control of spins
- *can this become too-ultrafast?*

Optics

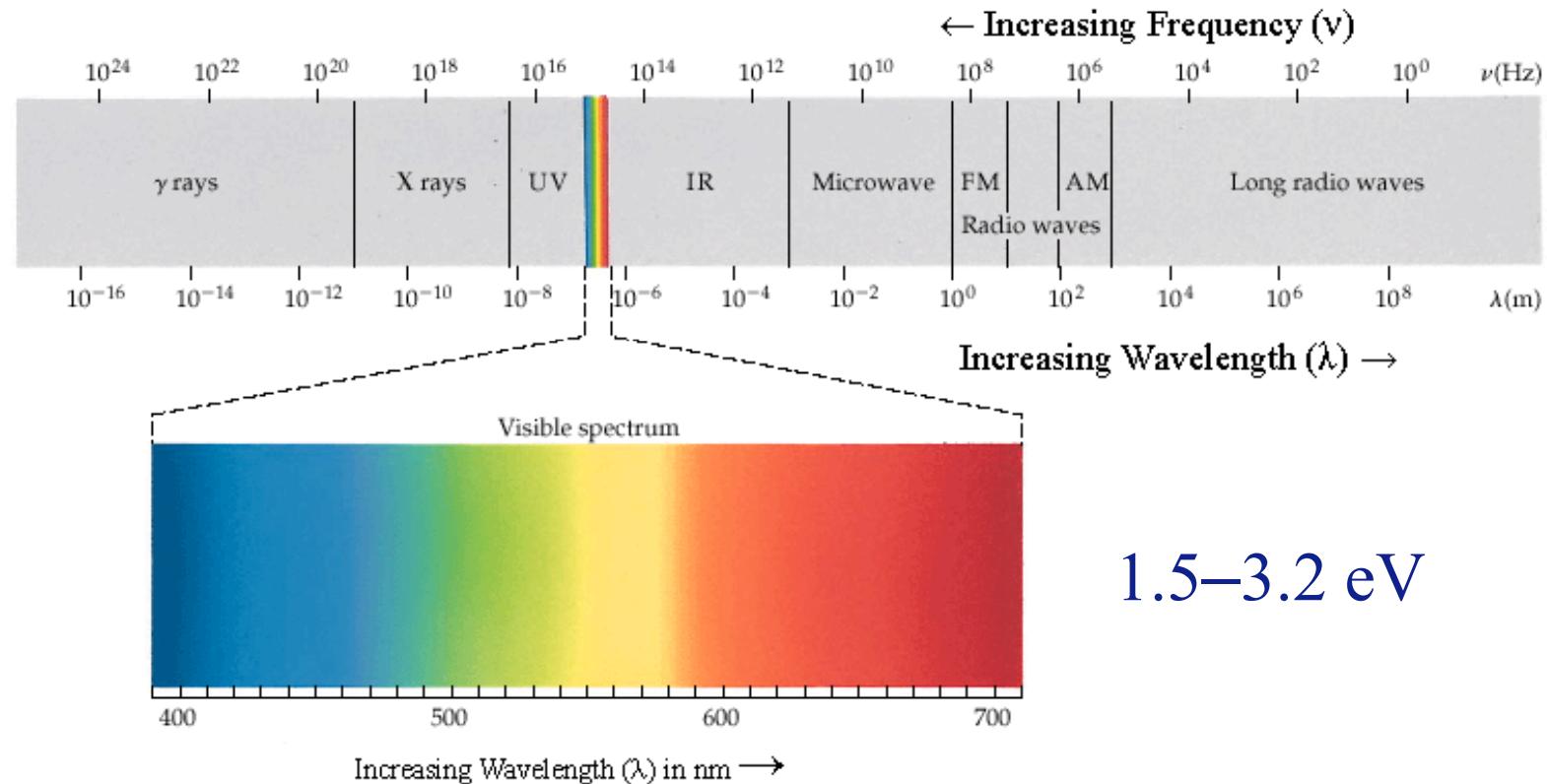


Why are certain wavelengths “visible”?

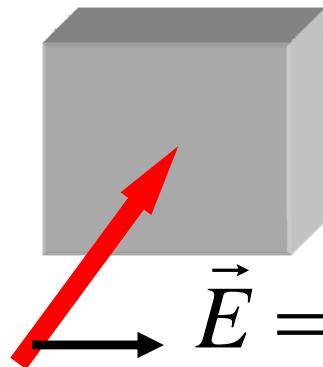
Transmission through water



Optics



1.5–3.2 eV



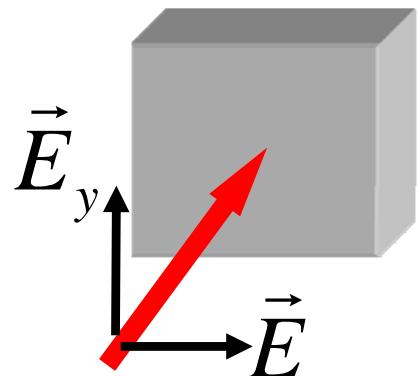
$$\vec{D} = \hat{\epsilon} \epsilon_0 \vec{E} \quad \text{or} \quad \vec{P} = \hat{\chi} \epsilon_0 \vec{E}$$

$$\vec{E} = \vec{E}_0 e^{i(\vec{k}\vec{r} - \omega t)} \quad \vec{D} = \epsilon_0 \vec{E} + \vec{P} \quad \hat{\epsilon} = 1 + \hat{\chi}$$

$$\epsilon = n^2$$

Anisotropic media

$$\hat{\mathcal{E}} = \begin{pmatrix} \mathcal{E}_{xx} & \mathcal{E}_{xy} & \mathcal{E}_{xz} \\ \mathcal{E}_{yx} & \mathcal{E}_{yy} & \mathcal{E}_{yz} \\ \mathcal{E}_{zx} & \mathcal{E}_{zy} & \mathcal{E}_{zz} \end{pmatrix}$$

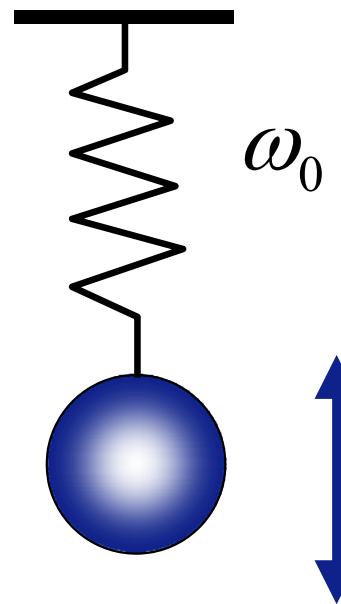
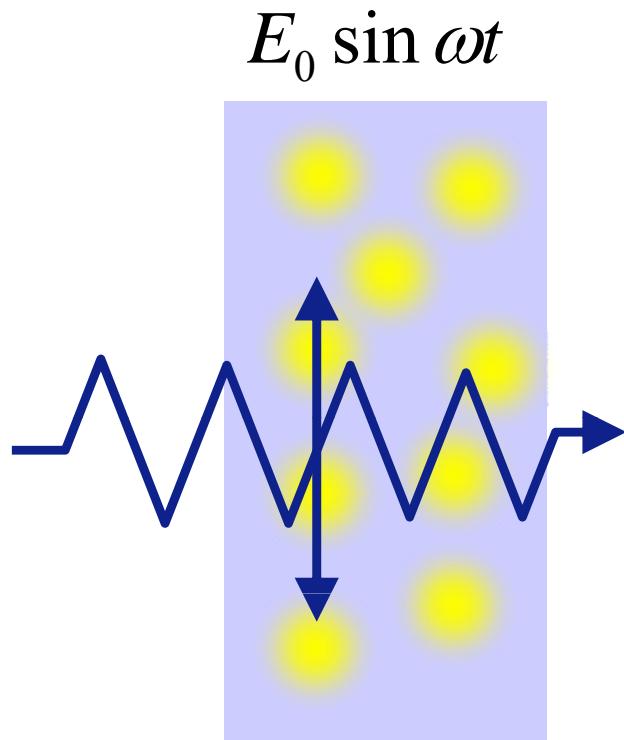


if $H = 0$

$$\hat{\mathcal{E}} = \begin{pmatrix} \mathcal{E}_{xx} & 0 & 0 \\ 0 & \mathcal{E}_{yy} & 0 \\ 0 & 0 & \mathcal{E}_{zz} \end{pmatrix}$$

i.e. one could chose such a coordinate system

Driven oscillator model

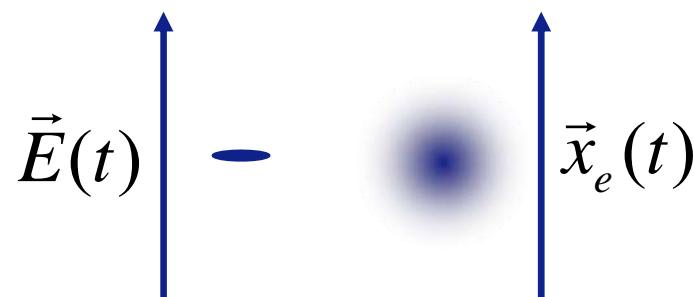


$$m \left(\frac{d^2 x}{dt^2} + \omega_0^2 x \right) = F$$

$$F = eE_0 \sin \omega t$$

$$\frac{d^2 x}{dt^2} + \omega_0^2 x = \frac{e}{m} E_0 \sin \omega t$$

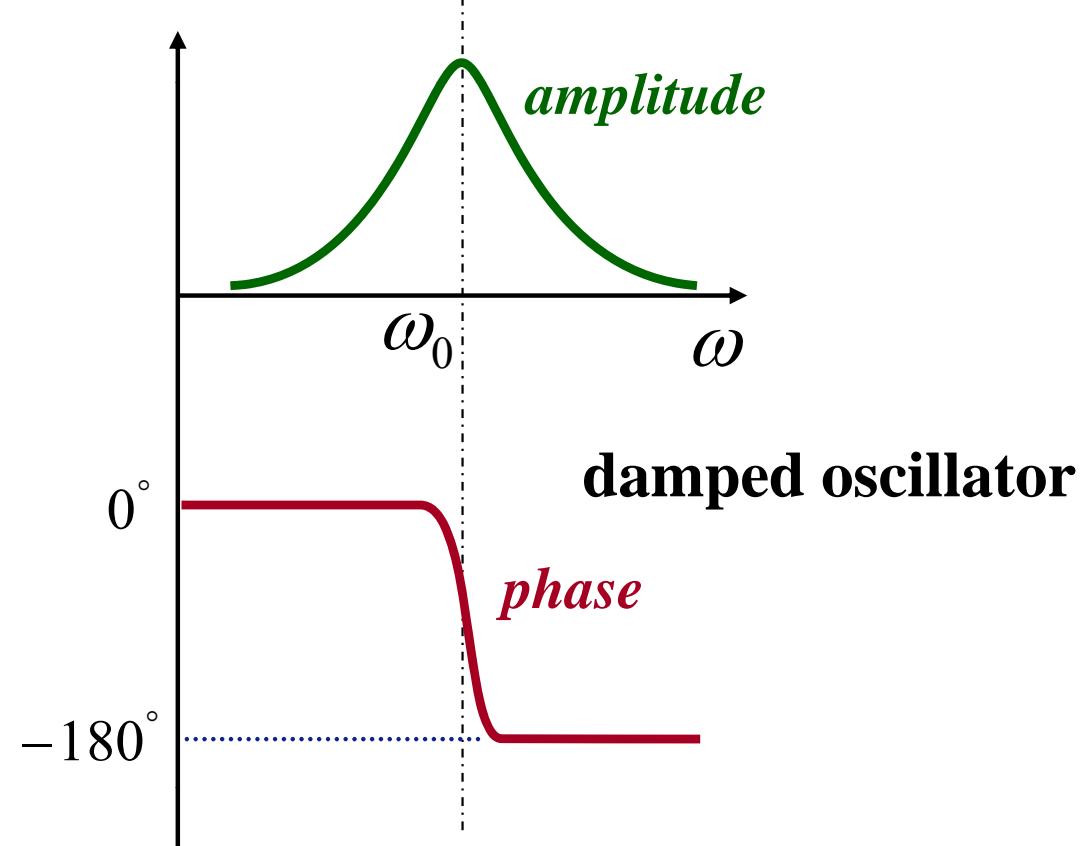
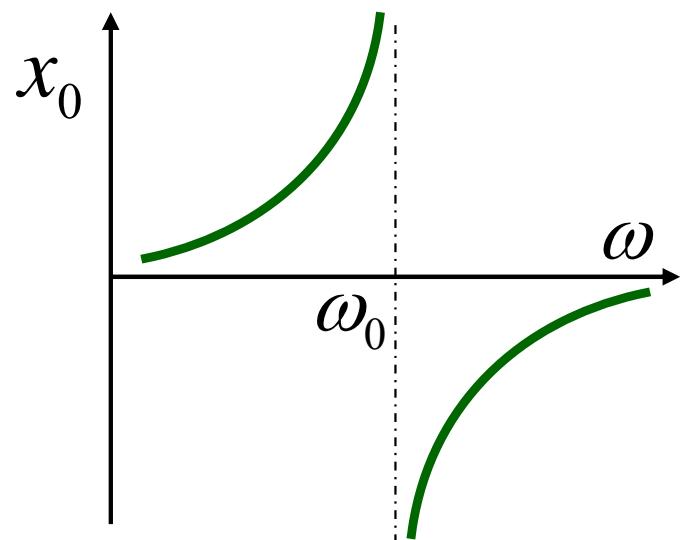
Driven oscillator model



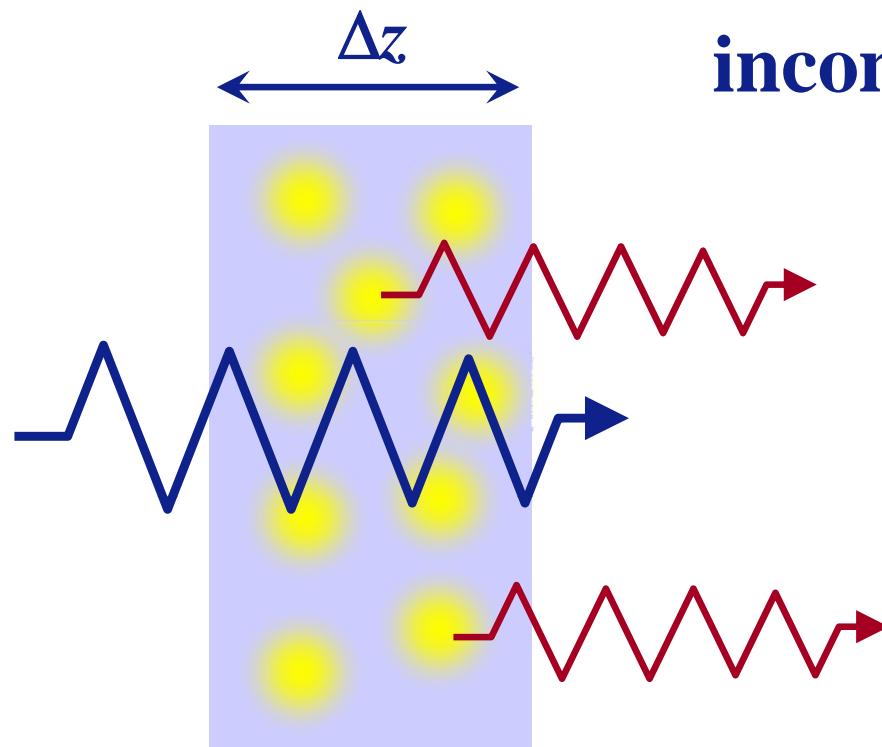
solution in the form:

$$x = x_0 \sin \omega t = \frac{eE_0}{m(\omega_0^2 - \omega^2)} \sin \omega t$$

amplitude depends on ω

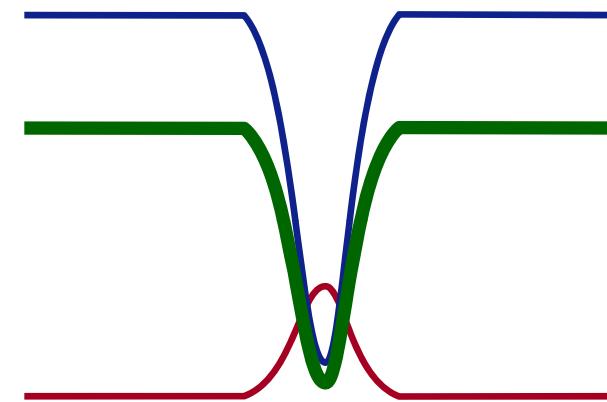


Sum of the two waves:

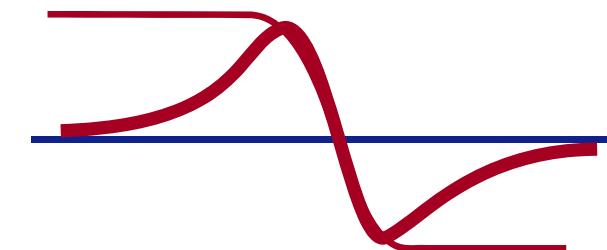


incoming + outgoing

amplitude

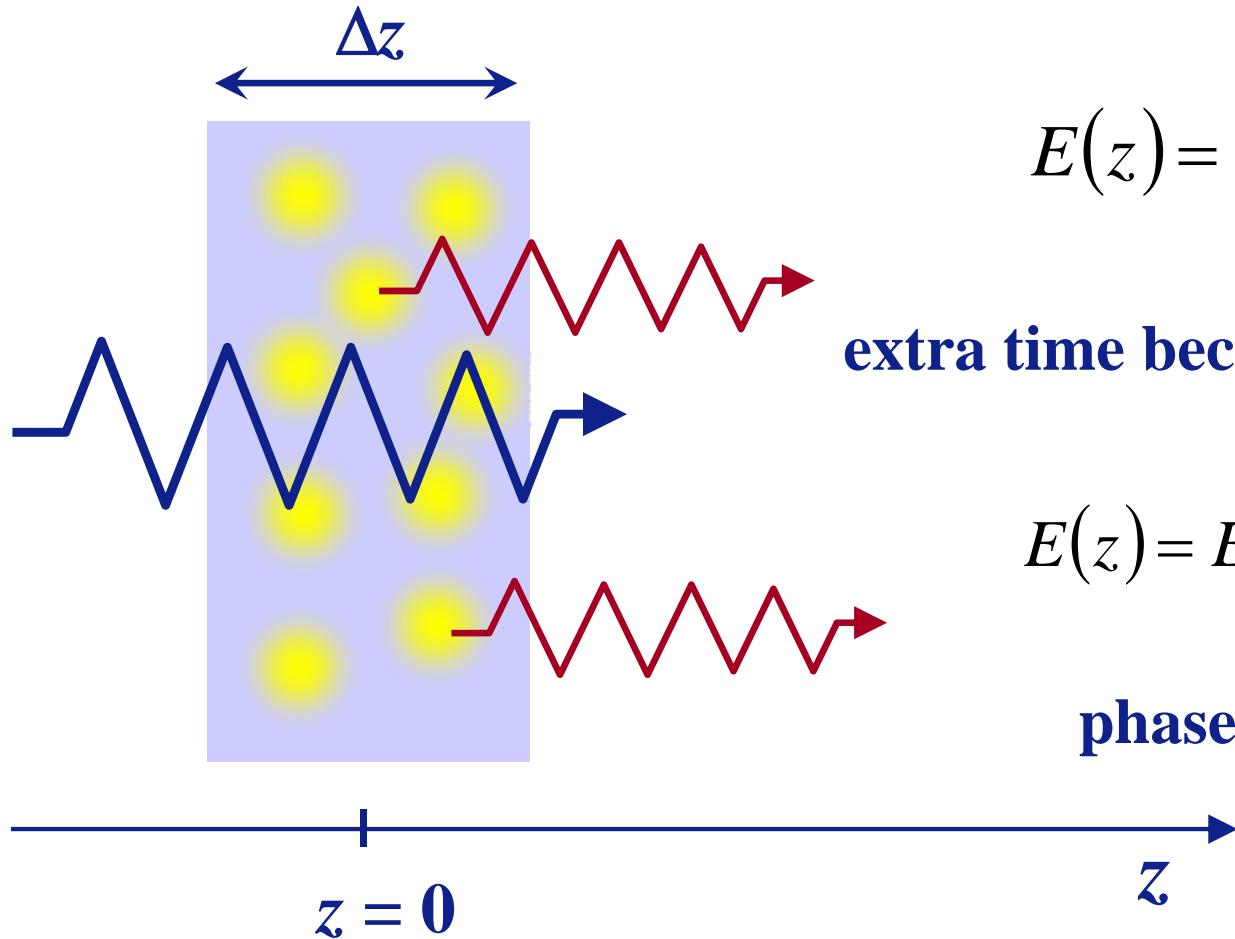


phase



$$E(z = 0) = E_0 \sin \omega t$$

Phase of the light after transmission



$$E(z) = E_0 \sin \omega \left(t - \frac{z}{c} \right)$$

extra time because of n : $(n-1) \frac{\Delta z}{c}$

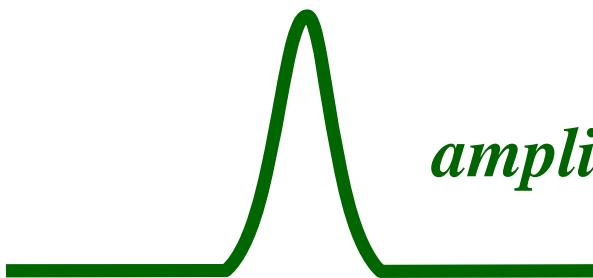
$$E(z) = E_0 \sin \omega \left(t - \frac{z}{c} - (n-1) \frac{\Delta z}{c} \right)$$

phase delay: $\omega(n-1) \frac{\Delta z}{c}$

$$E(z=0) = E_0 \sin \omega t$$

**thus if a phase delay occurs,
this is equivalent to the refractive index**

Absorption en refractive index

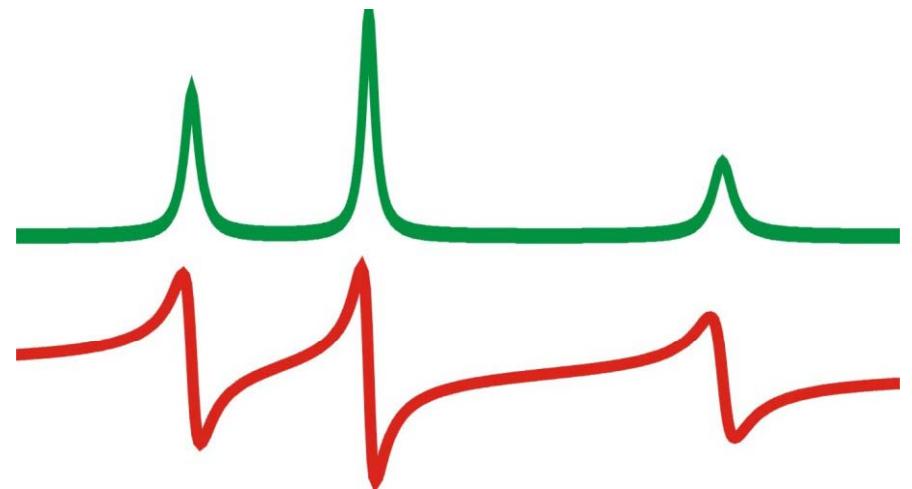


amplitude = absorption



phase = refractive index

multiple resonances:



$$\alpha = \frac{Ne^2}{2\epsilon_0 c_0 m_e} \frac{\gamma / 2}{(\omega_0 - \omega)^2 + (\gamma / 2)^2}$$

$$n - 1 = \frac{Ne^2}{4\epsilon_0 \omega m_e} \frac{\omega_0 - \omega}{(\omega_0 - \omega)^2 + (\gamma / 2)^2}$$

γ accounts for damping

Kramers-Kronig relations

$$\varepsilon = \varepsilon_1 + i\varepsilon_2$$

$$\tilde{n} = n + ik$$

$$\varepsilon_1 = n^2 - k^2$$

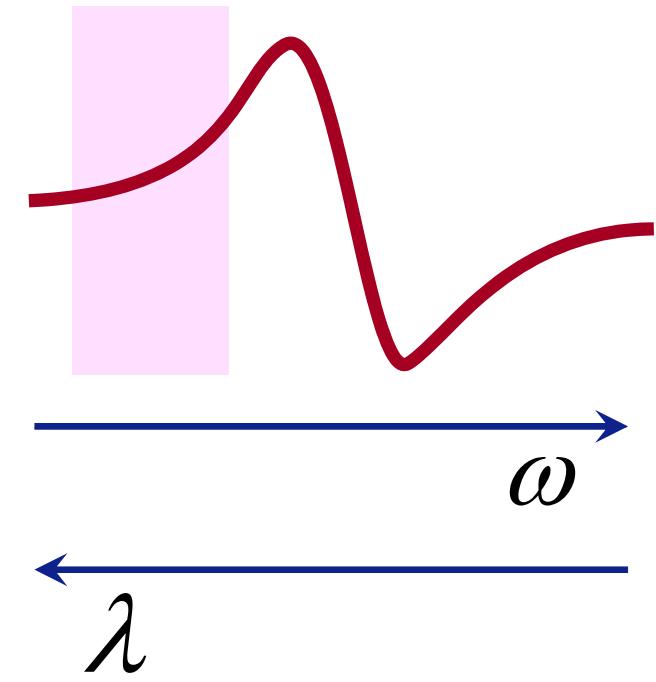
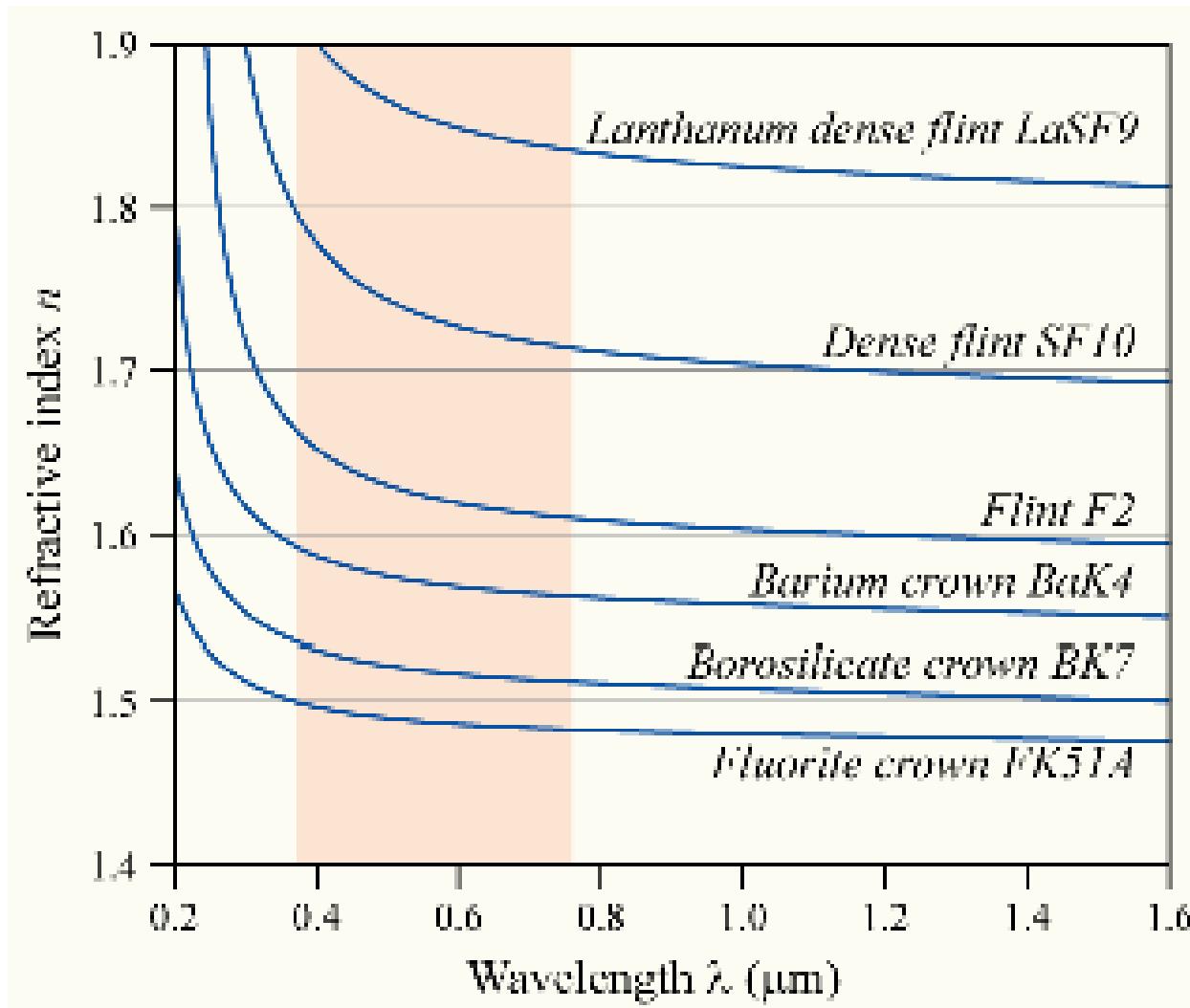
$$\varepsilon_2 = 2nk$$

$$\varepsilon_2(\omega) = \frac{2\omega}{\pi} \int_0^\infty \frac{1 - \varepsilon_1(u)}{u^2 - \omega^2} du$$

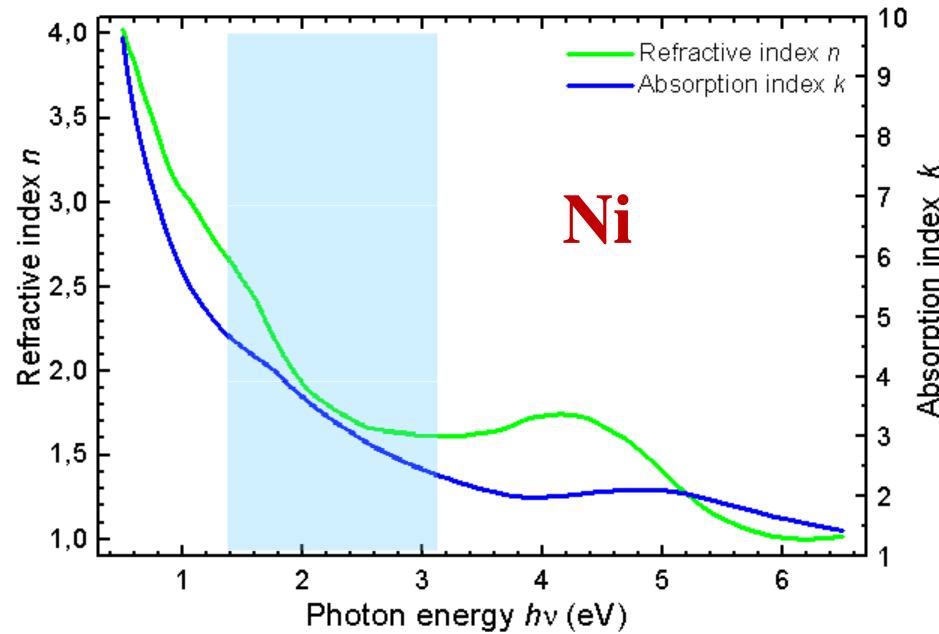
$$\varepsilon_1(\omega) = \frac{2}{\pi} \int_0^\infty \frac{u \varepsilon_2(u)}{u^2 - \omega^2} du$$

real and imaginary parts are not independent!

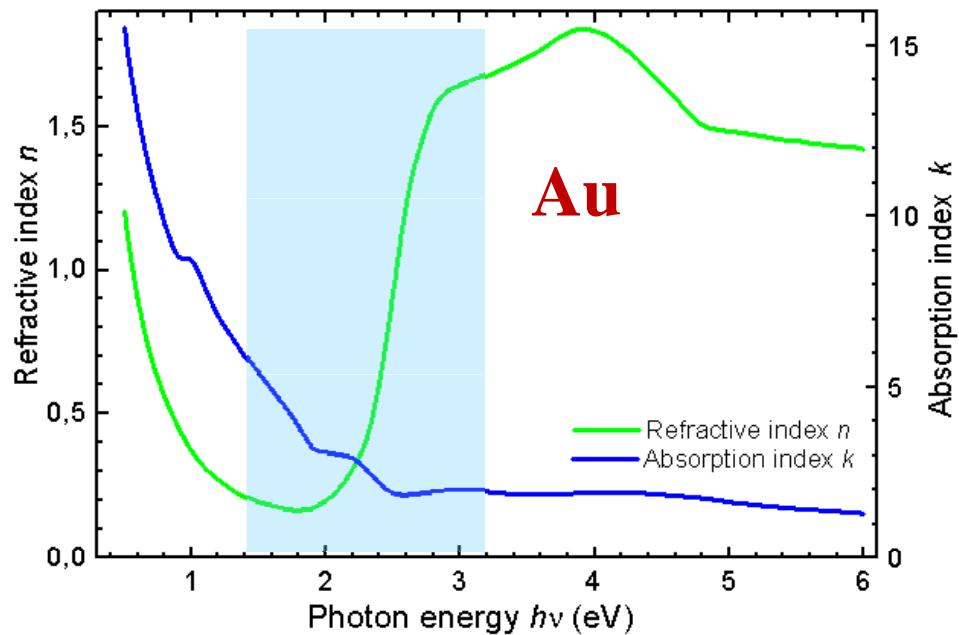
Dispersion of glass



Optical constants of metals



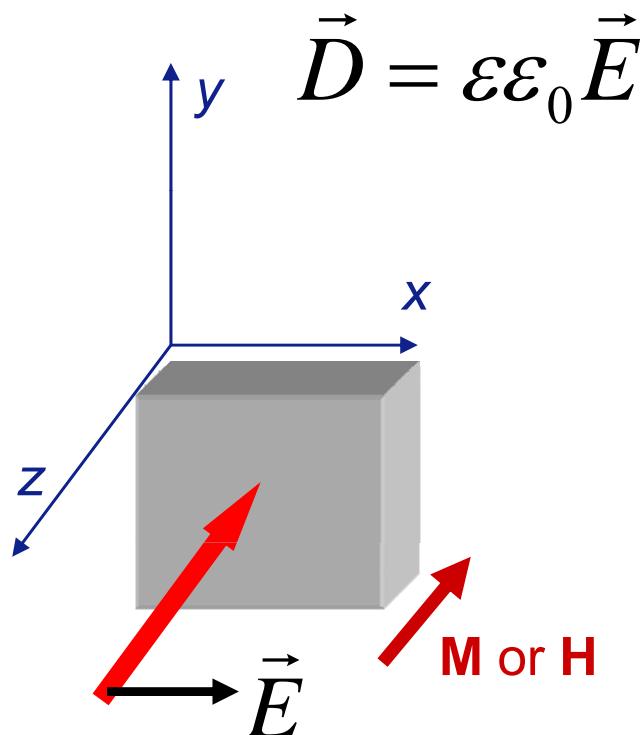
Ni



Au

Interaction of light with magnetic solids

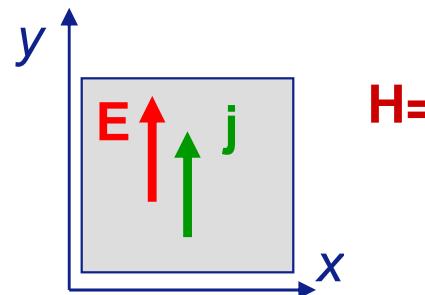
How does magnetic field (magnetization) modify dielectric tensor?



$$H=0 \quad \epsilon = \begin{bmatrix} \epsilon_{xx} & 0 & 0 \\ 0 & \epsilon_{yy} & 0 \\ 0 & 0 & \epsilon_{zz} \end{bmatrix} = n^2 \quad \text{if isotropic}$$

$$H \neq 0 \quad \epsilon = \begin{bmatrix} ? & ? & ? \\ ? & ? & ? \\ ? & ? & ? \end{bmatrix}$$

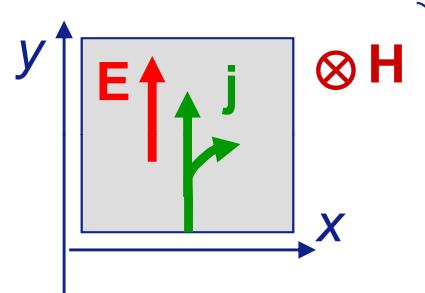
How does magnetic field modify conductivity?



$H=0$

$$\sigma = \begin{bmatrix} \sigma_{xx} & 0 & 0 \\ 0 & \sigma_{yy} & 0 \\ 0 & 0 & \sigma_{zz} \end{bmatrix}$$

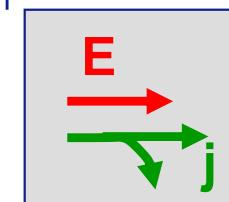
$$\vec{j} = \sigma \vec{E}$$



$\otimes H$

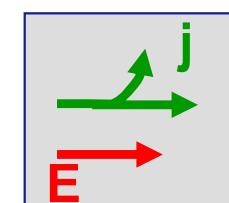
$$F_L = e[V \times H]$$

Lorentz force



$\otimes H$

$$E_y \rightarrow j_x$$
$$E_x \rightarrow j_y$$



$\bullet H$

$$\sigma = \begin{bmatrix} \sigma_{xx} & \sigma_{xy} & 0 \\ \sigma_{yx} & \sigma_{yy} & 0 \\ 0 & 0 & \sigma_{zz} \end{bmatrix}$$

Onsager principle: symmetry of kinetic coefficients

$$\mathbf{H=0} \quad \boldsymbol{\varepsilon} = \begin{bmatrix} \varepsilon_{xx} & 0 & 0 \\ 0 & \varepsilon_{yy} & 0 \\ 0 & 0 & \varepsilon_{zz} \end{bmatrix}$$

$$\mathbf{H \neq 0} \quad \boldsymbol{\varepsilon} = \begin{bmatrix} \varepsilon_{xx} & \varepsilon_{xy} & 0 \\ \varepsilon_{yx} & \varepsilon_{yy} & 0 \\ 0 & 0 & \varepsilon_{zz} \end{bmatrix}$$

$$X_i = S_{ij} f_j \quad \frac{\partial W}{\partial t} = f \frac{\partial X}{\partial t}$$

X - response
 f - stimulus
 W - energy



$$S_{ij} = S_{ji}$$

If S is a function of magnetic field

$$S_{ij}(H) = -S_{ji}(H) = S_{ji}(-H)$$

$$D_i = \varepsilon_{ij} E_j \quad \frac{\partial W}{\partial t} = E \frac{\partial D}{\partial t}$$

Onsager principle
is applicable to ε

$$\varepsilon_{ij}(H) = -\varepsilon_{ji}(H) = \varepsilon_{ji}(-H)$$

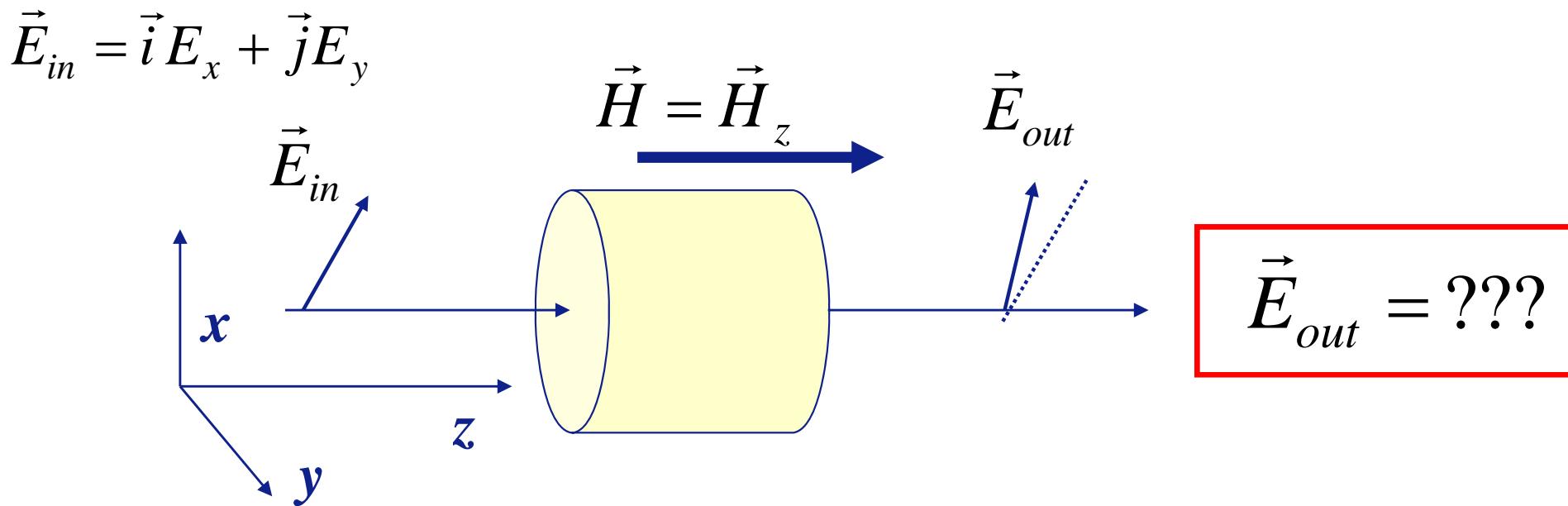
in non-absorbing media $\varepsilon_{ij} = \varepsilon_{ji}^*$

Landau & Lifshitz, Theoretical Physics, vv. 5 and 8.

Faraday effect – 1

Isotropic medium in a magnetic field:

$$\hat{\epsilon} = \begin{pmatrix} \epsilon_0 & -i\epsilon_{xy} & 0 \\ i\epsilon_{xy} & \epsilon_0 & 0 \\ 0 & 0 & \epsilon_0 + \epsilon_{zz} \end{pmatrix} \quad \epsilon_{xy} \propto M_z \quad \epsilon_{zz} \propto M_z^2$$



Faraday effect – 2

To find the eigenvalues of the problem:

$$\vec{D} = \hat{\epsilon} \vec{E} = \begin{pmatrix} \epsilon_0 & -i\epsilon_{xy} & 0 \\ i\epsilon_{xy} & \epsilon_0 & 0 \\ 0 & 0 & \epsilon_0 \end{pmatrix} \begin{pmatrix} E_x \\ E_y \\ 0 \end{pmatrix} = n^2 \vec{E} \quad \begin{pmatrix} \epsilon_0 & -i\epsilon_{xy} \\ i\epsilon_{xy} & \epsilon_0 \end{pmatrix} \begin{pmatrix} E_x \\ E_y \end{pmatrix} = n^2 \begin{pmatrix} E_x \\ E_y \end{pmatrix}$$

$$\begin{vmatrix} \epsilon_0 - n^2 & -i\epsilon_{xy} \\ i\epsilon_{xy} & \epsilon_0 - n^2 \end{vmatrix} = 0 \quad \Rightarrow \quad n^2 = \epsilon_0 \pm \epsilon_{xy}$$
$$n \cong \sqrt{\epsilon_0} \pm \frac{1}{2} \frac{\epsilon_{xy}}{\epsilon_0} \quad \epsilon_{xy} \sim 10^{-3} - 10^{-4}$$

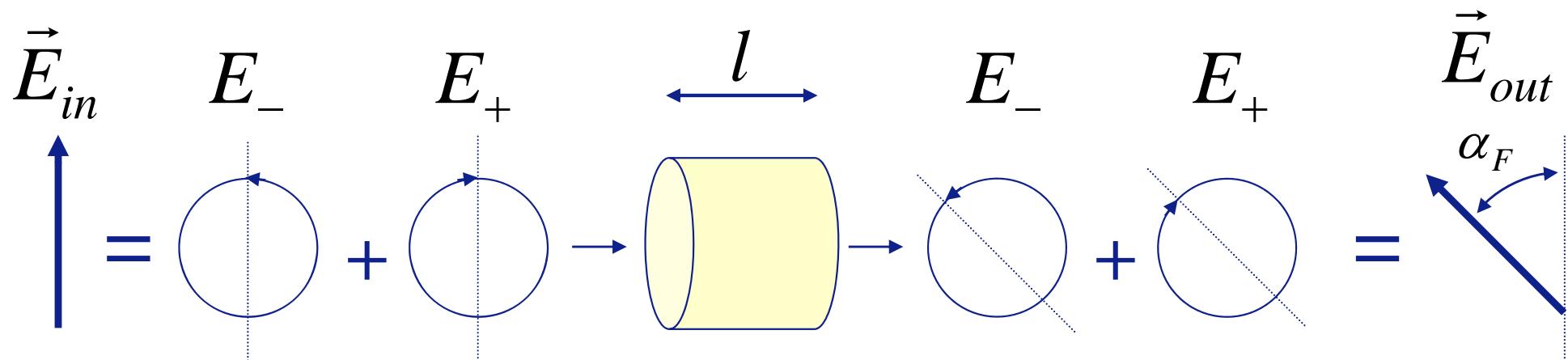
eigenmodes $E_x = \pm iE_y$, or $\begin{pmatrix} 1 \\ i \end{pmatrix}$ and $\begin{pmatrix} 1 \\ -i \end{pmatrix}$

Faraday effect – 3

Two circularly polarized waves with different refractive indices:

$$E_x = \pm i E_y$$

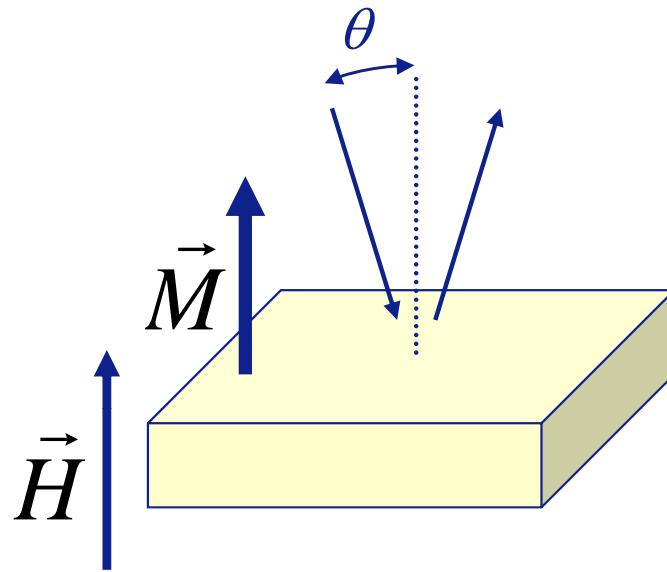
$$n_{\pm} \cong \sqrt{\epsilon_0} \pm \frac{1}{2} \frac{\epsilon_{xy}}{\epsilon_0}$$



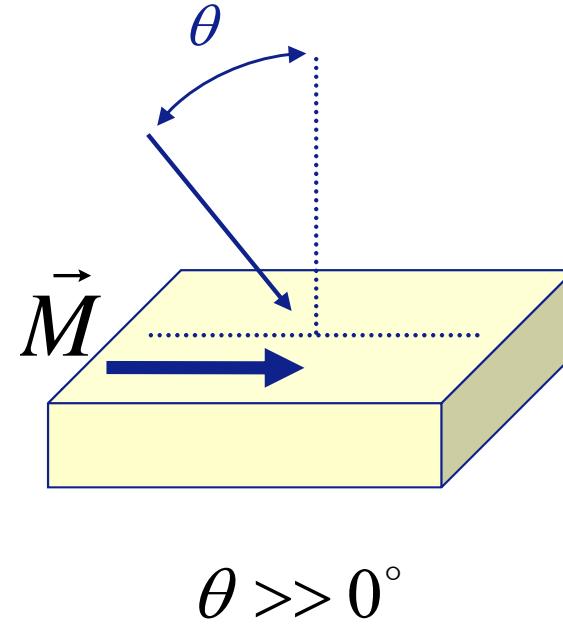
$$\text{Faraday rotation: } \alpha_F = \frac{2\pi l}{\lambda} \frac{\epsilon_{xy}}{\epsilon_0}$$

M. Faraday, *On the magnetization of light and the illumination of magnetic lines of force*, Phil. Trans. R. Soc. Lond. **136**, 104 (1846).

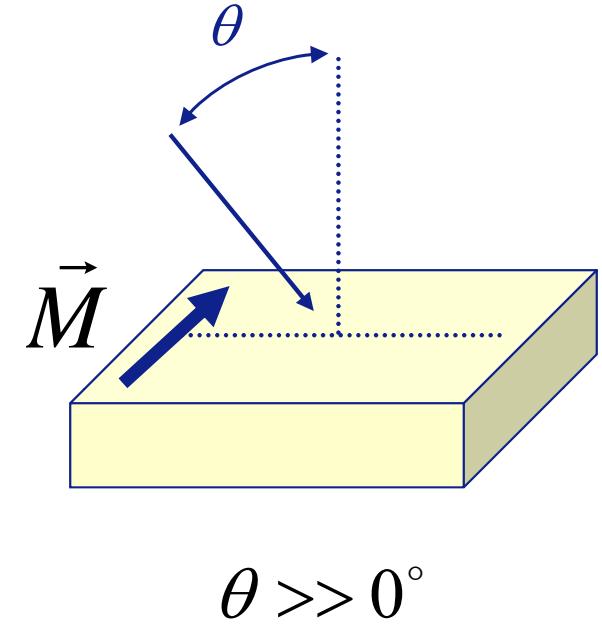
Kerr effect: various geometries



polar



longitudinal



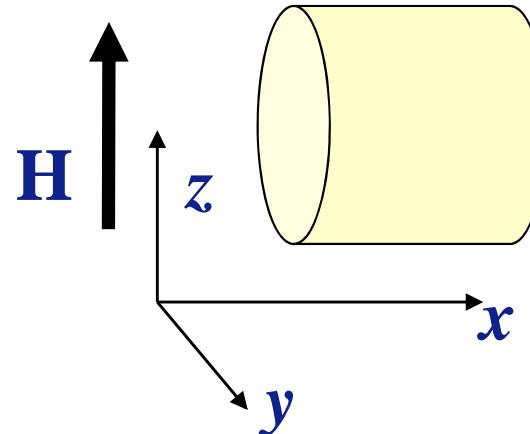
transverse

Magnetic linear birefringence

light propagates along x axis, so that

$$\vec{E}_{in} = \vec{j}E_y + \vec{k}E_z$$

$$\hat{\epsilon} = \begin{pmatrix} \epsilon_0 & -i\epsilon_{xy} & 0 \\ i\epsilon_{xy} & \boxed{\epsilon_0} & 0 \\ 0 & 0 & \epsilon_0 + \epsilon_{zz} \end{pmatrix}$$



Eigenvalues $\epsilon_0, \epsilon_0 + \epsilon_{zz}, \epsilon_{zz} \propto M^2$

Eigenmodes \vec{E}_y, \vec{E}_z

Theory of the Faraday and Kerr Effects in Ferromagnetics*

PETROS N. ARGYRES†

Department of Physics, University of California, Berkeley, California

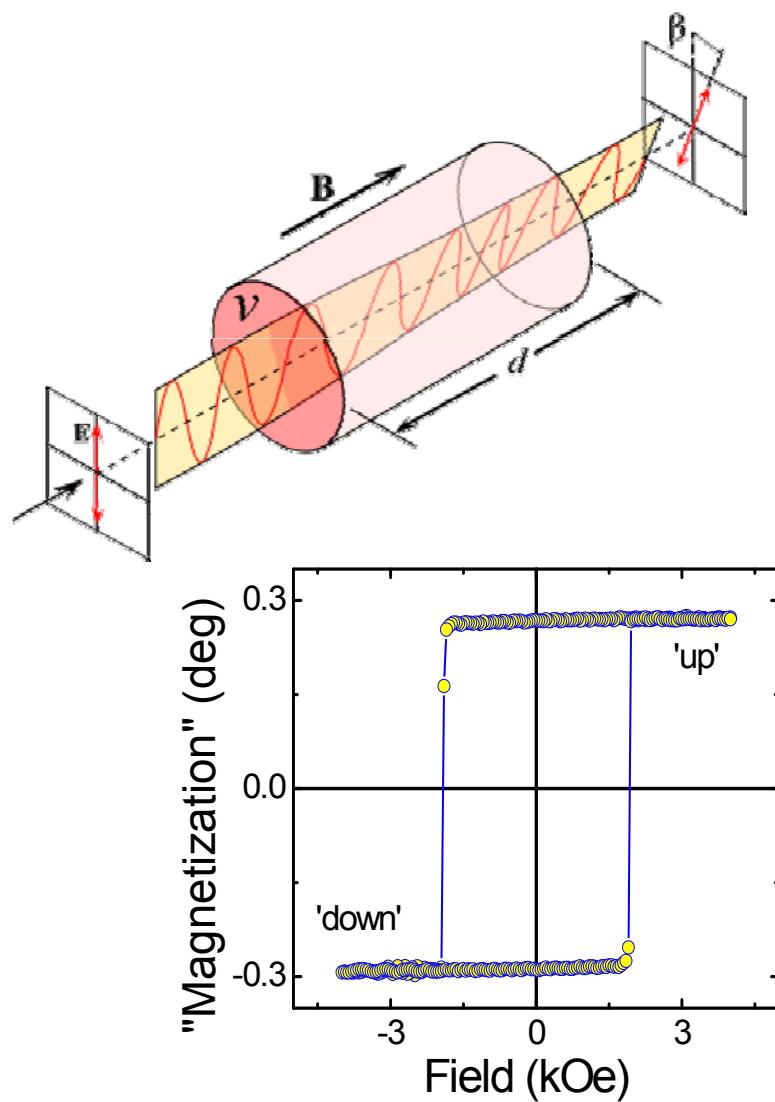
(Received August 20, 1954)

Both the Faraday and (magneto-optic, polar) Kerr effects in ferromagnetics are treated on the basis of the band theory of metals. The spin-orbit interaction gives the electron wave functions such left-right asymmetry that the "magnetic" electrons, under the action of a plane polarized light wave, produce an average current perpendicular to the plane of polarization.

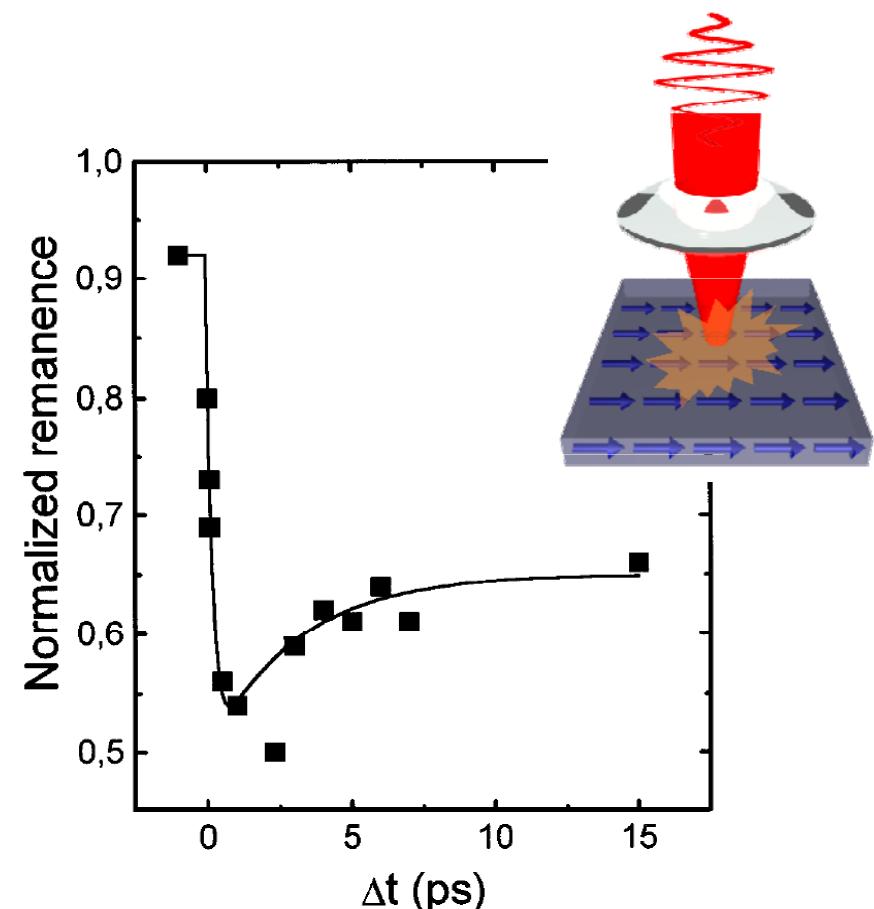
spins → orbits → light wave

exchange + spin-orbit

What could be measured?



hysteresis



Beaurepaire et al, PRL 76, 4250 (1996).

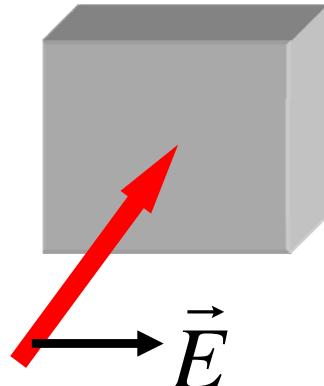
dynamics

Outline of the lecture

- Light as a probe
 - linear magneto-optics
 - nonlinear (magneto-)optics
- Example: all-optical FMR
- Light as an excitation
 - classification of effects
 - basics of opto-magnetism
 - coherent control
 - local control of spins
- *can this become too-ultrafast?*

Electromagnetic wave equation and source term

$$\frac{\partial^2 \vec{E}}{\partial t^2} - c^2 \cdot \nabla^2 \vec{E} = \vec{S} \quad \vec{S} = \frac{\partial^2 \vec{P}}{\partial t^2} + \nabla \times \frac{\partial \vec{M}}{\partial t} - \nabla \frac{\partial^2 \hat{Q}}{\partial t^2}$$



$$\Phi = - \left(\underbrace{\chi^{(1,d)} E^\omega E^\omega}_{\dots} + \underbrace{\chi^{(1,m)} E^\omega H^\omega}_{\dots} + \underbrace{\chi^{(1,d)} E^\omega \nabla E^\omega}_{\dots} \right) \\ \times \left(1 + \alpha E^\omega + \beta H^\omega + \gamma \nabla E^\omega + \dots \right),$$

or

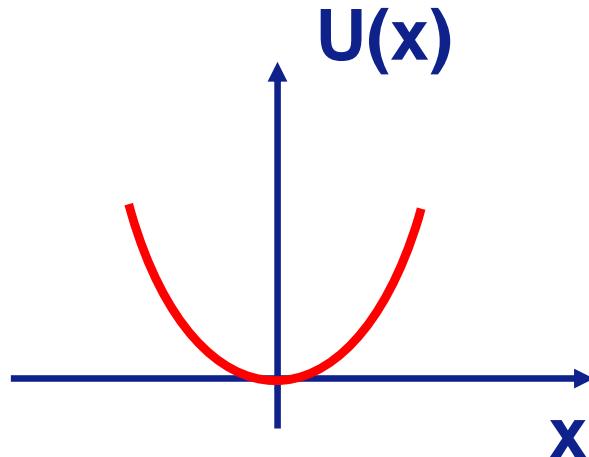
$$P^{\omega\dots} = \underbrace{\chi^{(1,d)} E^\omega}_{\dots} + \underbrace{\chi^{(1,m)} H^\omega}_{\dots} + \underbrace{\chi^{(1,q)} \nabla E^\omega}_{\dots} + \\ + \underbrace{\chi^{(2,d)} E^\omega E^\omega}_{\dots} + \underbrace{\chi^{(2,m)} E^\omega H^\omega}_{\dots} + \underbrace{\chi^{(2,q)} E^\omega \nabla E^\omega}_{\dots} \dots$$

electric dipole approximation

magnetic dipole **electric quadrupole**

Harmonic oscillator

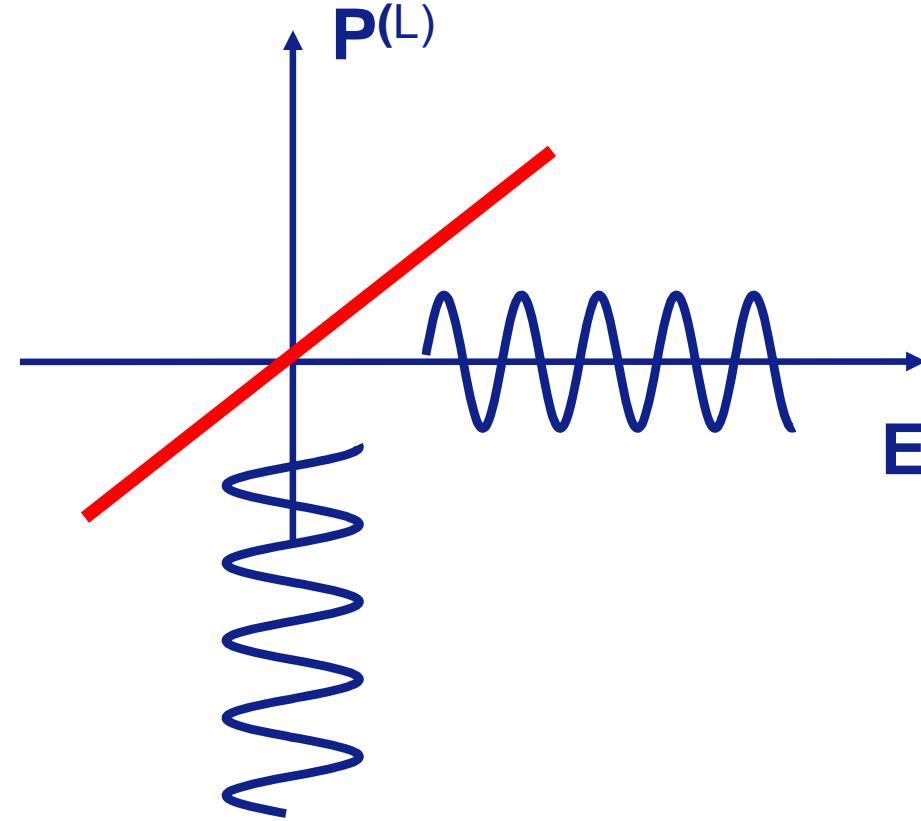
Linear response



$$U(x) = k_1 x^2$$

$$F_{el} = \frac{dU(x)}{dx} = kx$$

$$m \frac{d^2x}{dt^2} + F_{el} = eE$$



$$P_i^{(L)}(\omega) = \chi_{ij}(\omega) E_j(\omega)$$

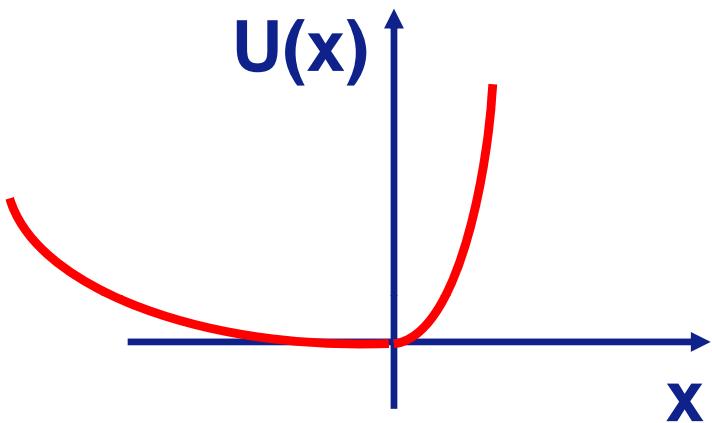
Linear vs nonlinear optics?

Linearity in optics:

- Properties of a medium ***do not depend*** on light intensity
- Principle of superposition ***holds***
- Frequency of light ***is not*** altered by its passage through the medium
- Light ***does not interact*** with light. A control of light by light is impossible.

Nonlinear oscillator

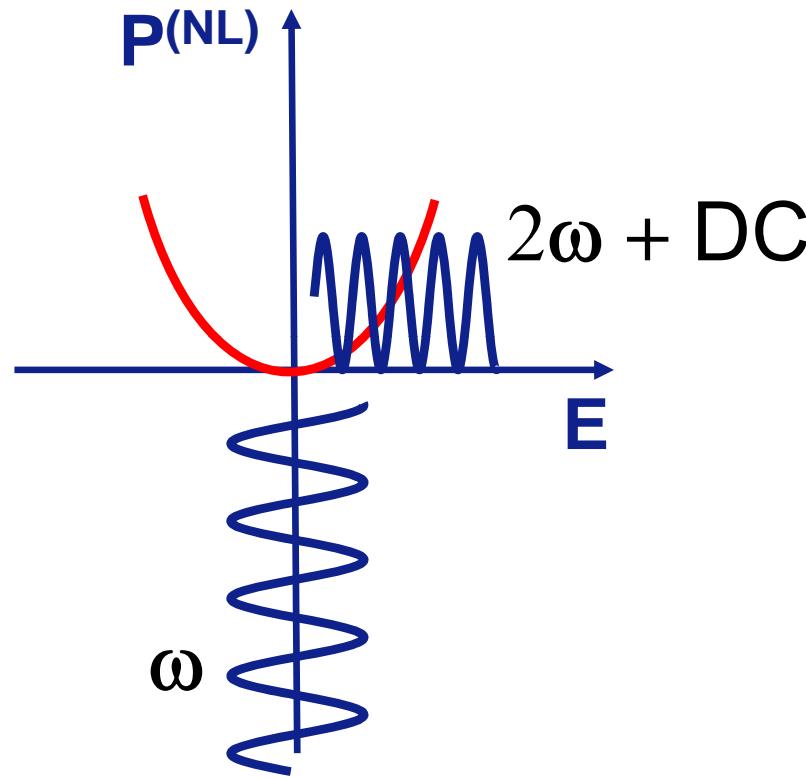
Anharmonism



$$U(x) = k_1 x^2 + k_2 x^3$$

$$F_{el} = \frac{dU(x)}{dx} = 2k_1 x + 3k_2 x^2$$

$$P = P^{(L)} + P^{(NL)}$$



$$P_i^{(NL)}(2\omega) = \chi_{ijk}(2\omega)E_j(\omega)E_k(\omega)$$

$$P_i^{(NL)}(0) = \chi_{ijk}(0)E_j(\omega)E_k(\omega)$$

Nonlinear polarization and symmetry

$$P(\omega) \propto \chi^{(1)} \cdot E^\omega + \chi^{(2)} \cdot E^\omega E^\omega + \dots$$

(electric dipole approximation)

$$e^{i\omega t} \quad e^{i\omega t} \cdot e^{i\omega t} = e^{i2\omega t}$$

second-harmonic generation

$$e^{i\omega t} \cdot e^{-i\omega t} = e^{i \cdot 0 t}$$

optical rectification

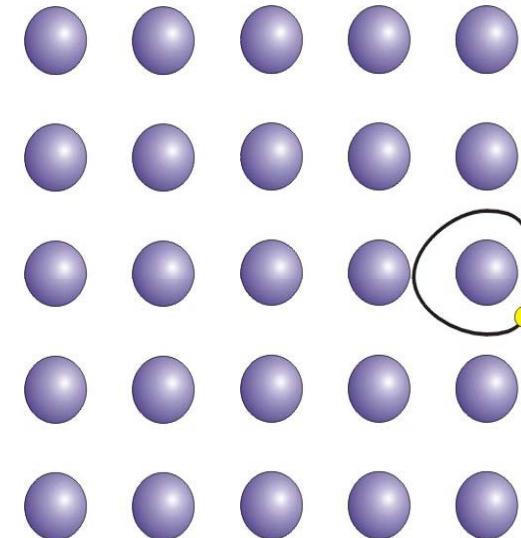
inversion symmetry:

$$P(2\omega) \propto \cancel{\chi^{(2)}} \cdot E^\omega E^\omega$$

-1 -1 -1

$$P(2\omega) = 0$$

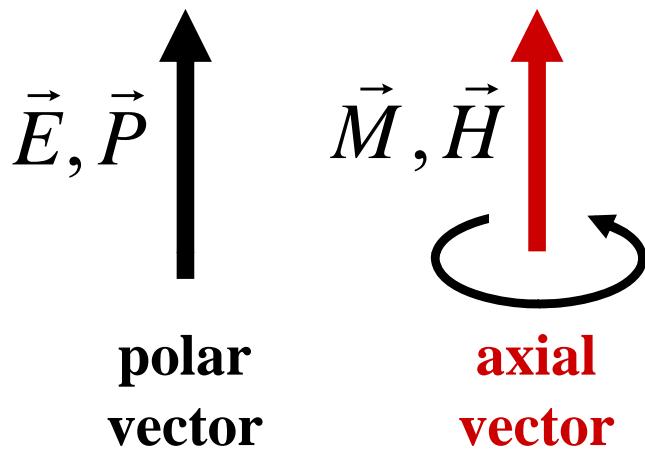
except at surface/interface



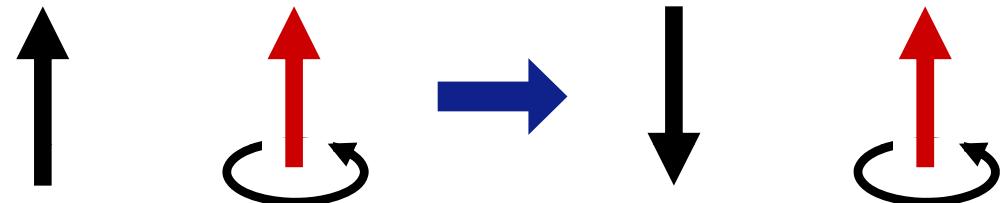
Magnetization-sensitive SHG

$$P_i^{(2\omega)} = \chi_{ijk} E_j^{(\omega)} E_k^{(\omega)} = (\chi_{ijk}^{(cr)} \pm \chi_{ijk}^{(m)}) E_j^{(\omega)} E_k^{(\omega)}$$

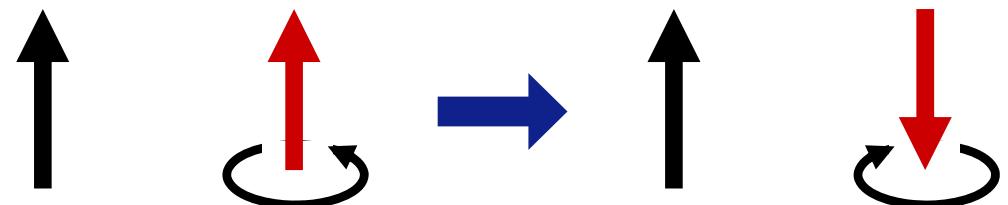
crystallographic
±
magnetic



space inversion:



time reversal:

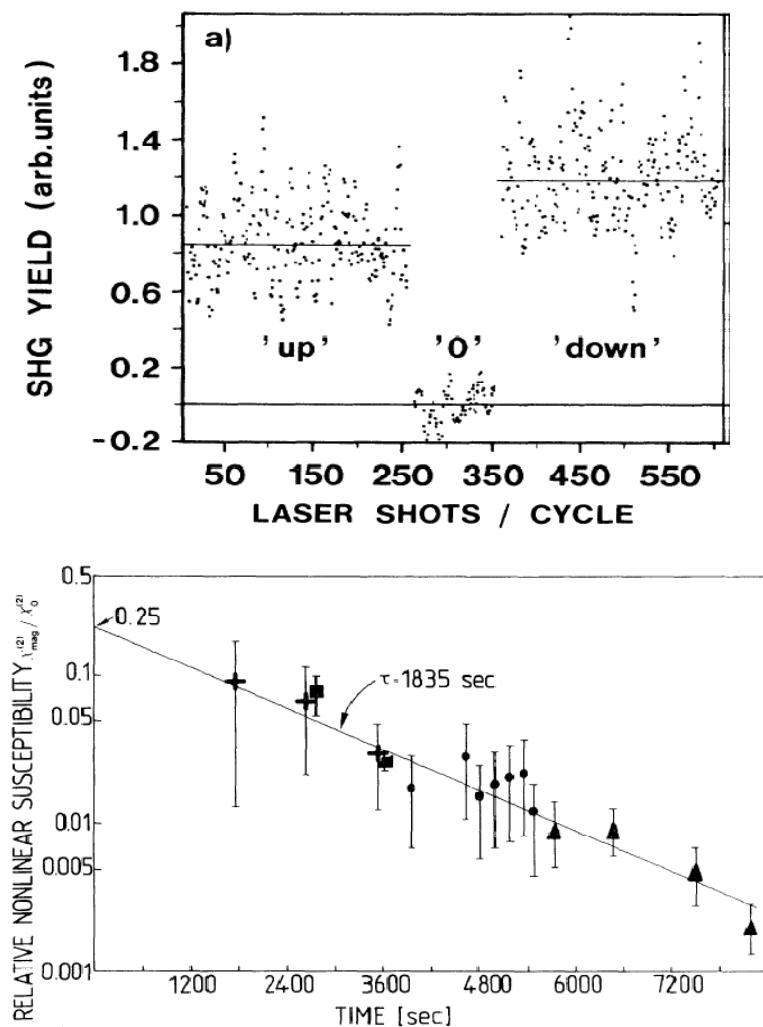


R.R. Birss, *Symmetry and Magnetism*
(North-Holland, Amsterdam, 1966).

A. Kirilyuk and Th. Rasing,
J. Opt. Soc. Am. B **22**, 148 (2005)

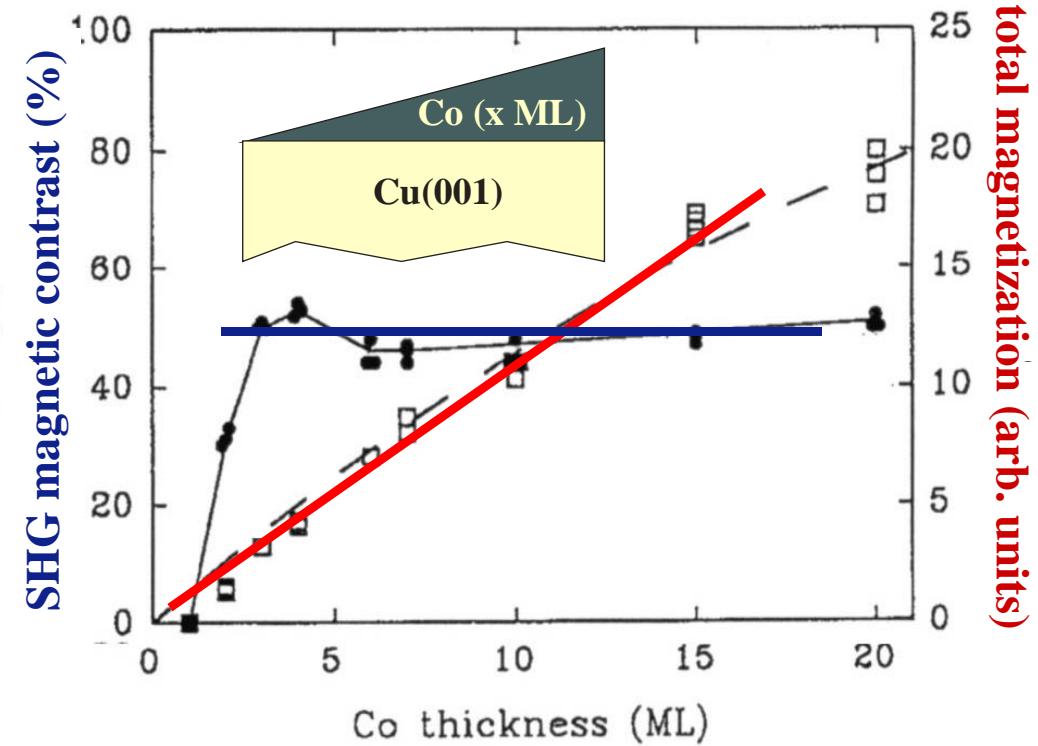
Example: surface/interface sensitivity

Fe(110) surface



Reif *et al*, Phys. Rev. Lett. **67**, 2878 (1991)

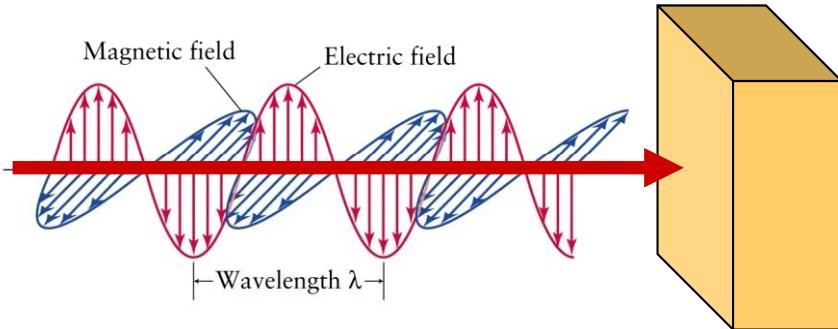
ultrathin Co/Cu(001) films



at 4 ML, both interfaces are formed

Phys. Rev. Lett. **74**, 1462 (1995);
J. Phys. D – Appl. Phys. **35**, R189 (2002)

General phenomenology



$$\vec{P} \propto \chi^{(1)} \cdot \vec{E} + \chi^{(2)} : \vec{E}\vec{E} + \chi^{(3)} : \vec{E}\vec{E}\vec{E} + \dots$$



$$\vec{P}^{(nl)} \propto \chi^{eee} : \vec{E}\vec{E} + \chi^{eem} : \vec{E}\vec{H} + \chi^{emm} : \vec{H}\vec{H} + o\left(\vec{E}, \vec{H}\right)^3$$

$$\vec{M}^{(nl)} \propto \chi^{mee} : \vec{E}\vec{E} + \chi^{mem} : \vec{E}\vec{H} + \chi^{mmm} : \vec{H}\vec{H} + o\left(\vec{E}, \vec{H}\right)^3$$

$$\hat{Q}^{(nl)} \propto \chi^{qee} : \vec{E}\vec{E} + \chi^{qem} : \vec{E}\vec{H} + \chi^{qmm} : \vec{H}\vec{H} + o\left(\vec{E}, \vec{H}\right)^3$$

Source term:

$$\mathbf{S} = \mu_0 \frac{\partial^2 \mathbf{P}}{\partial t^2} + \mu_0 \left(\nabla \times \frac{\partial \mathbf{M}}{\partial t} \right) - \mu_0 \left(\nabla \frac{\partial^2 \hat{Q}}{\partial t^2} \right)$$

**sum- and difference frequency generation,
including SHG and optical rectification**

Outline of the lecture

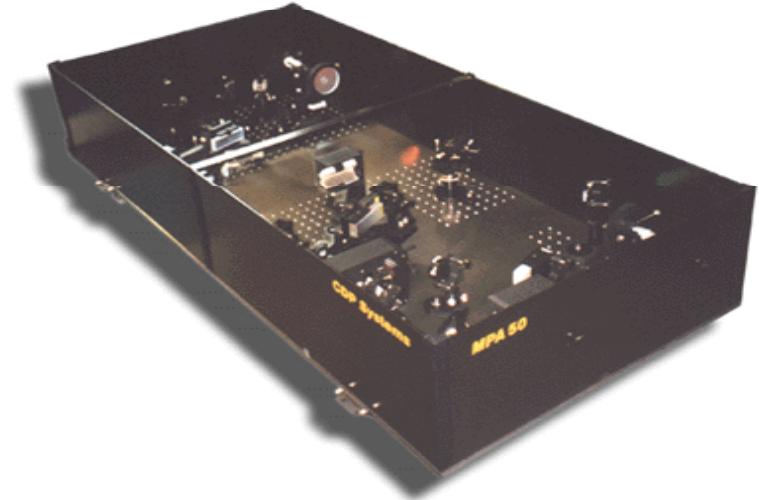
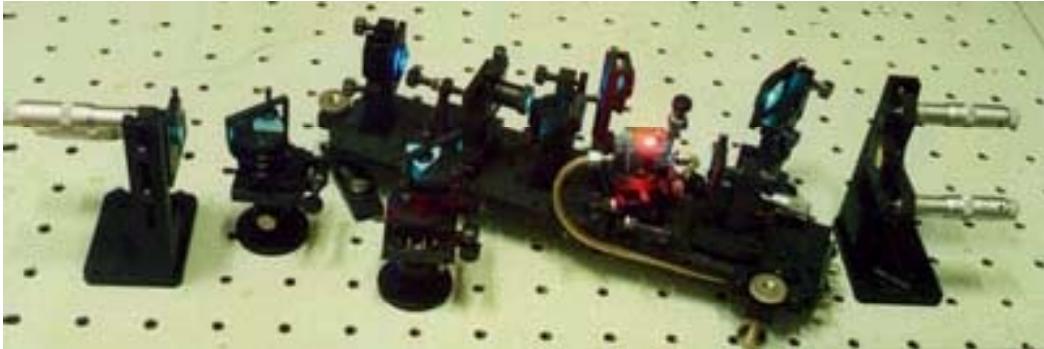
- Light as a probe ✓
 - linear magneto-optics ✓
 - nonlinear (magneto-)optics ✓
- Example: all-optical FMR
- Light as an excitation
 - classification of effects
 - basics of opto-magnetism
 - coherent control
 - local control of spins
- *can this become too-ultrafast?*

Outline of the lecture

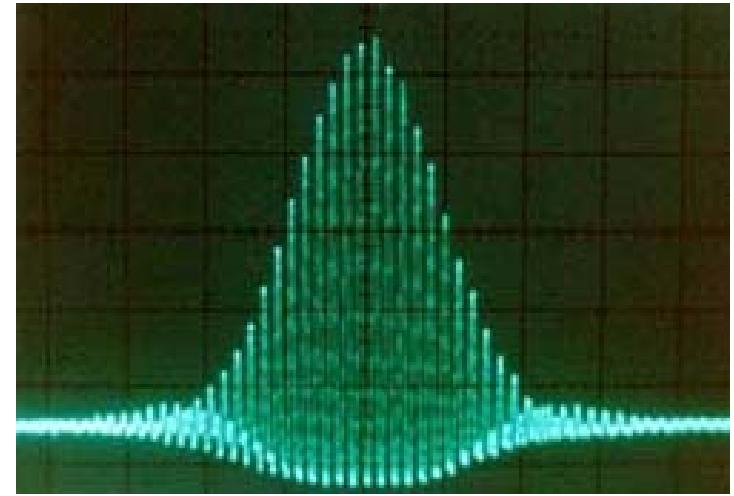
- Light as a probe
 - linear magneto-optics
 - nonlinear (magneto-)optics
- Example: all-optical FMR
- Light as an excitation
 - classification of effects
 - basics of opto-magnetism
 - coherent control
 - local control of spins
- *can this become too-ultrafast?*

Experimental know-how: time-resolved pump-probe setup

What you need: a femtosecond laser



	Model TISSA20	Model TISSA50	Model TISSA100
Pump Power ¹⁾	3-5 W	3-7 W	5-10 W
Output Power at 800 nm	150 - 250 mW	150-500 mW	>10% efficiency
Pulse Duration ²⁾	<20 fs ³⁾	< 50 fs	<100 fs
Tuning Range	800 ± 20 nm	740 - 950 nm ⁴⁾	720 - 980 nm ⁴⁾
Repetition Rate	70 - 140 MHz		



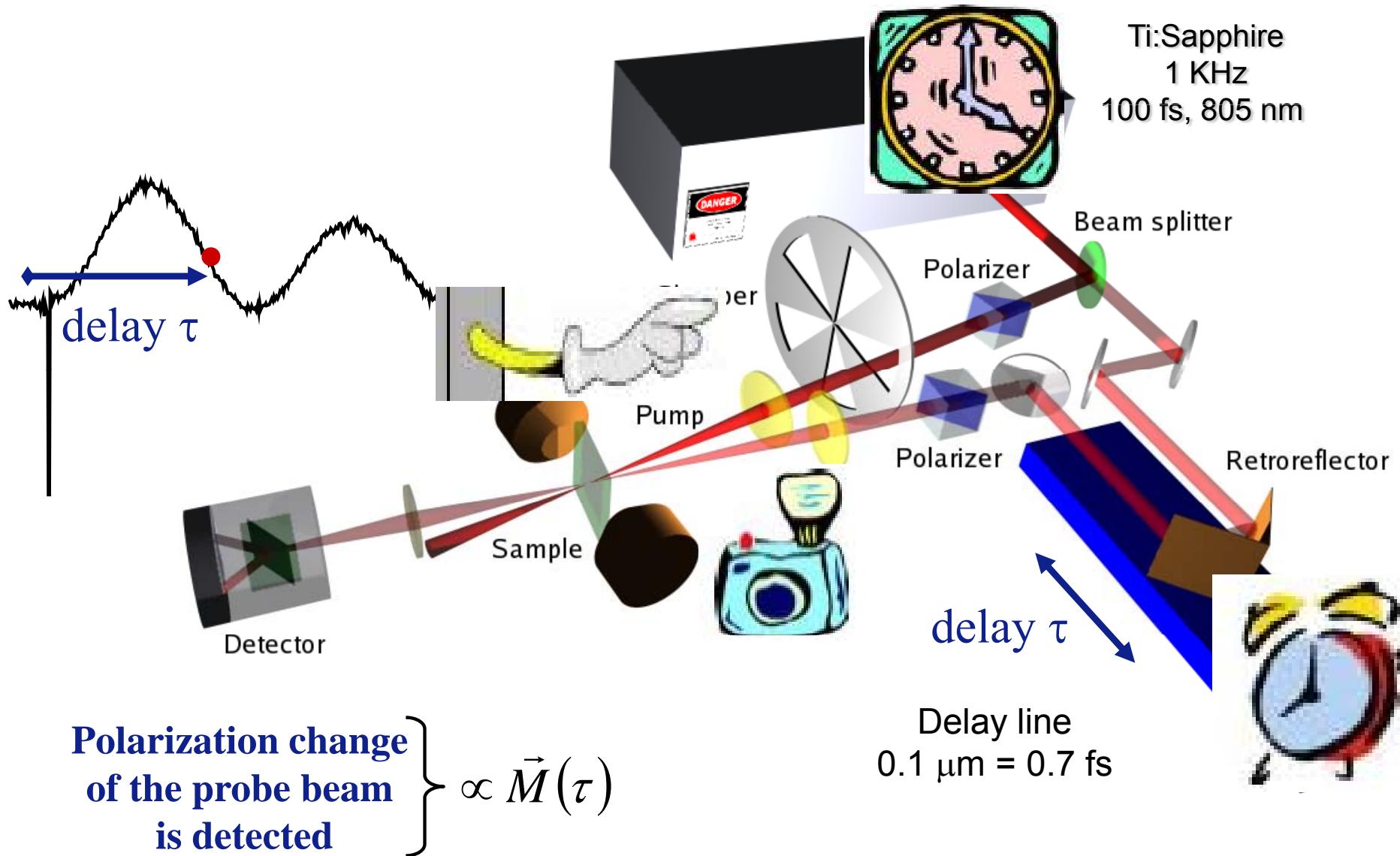
*Interferometric autocorrelation function
of 16 fs pulse obtained with external
group velocity dispersion compensation*

you have some choice!

pump & probe technique

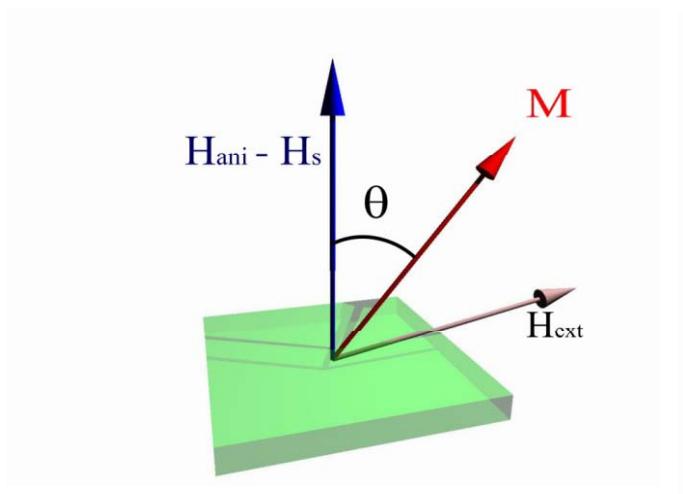


Stroboscopic magneto-optical pump-probe measurements

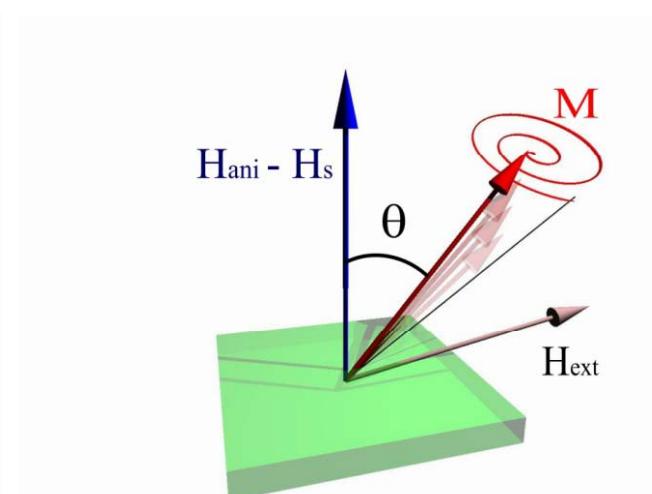
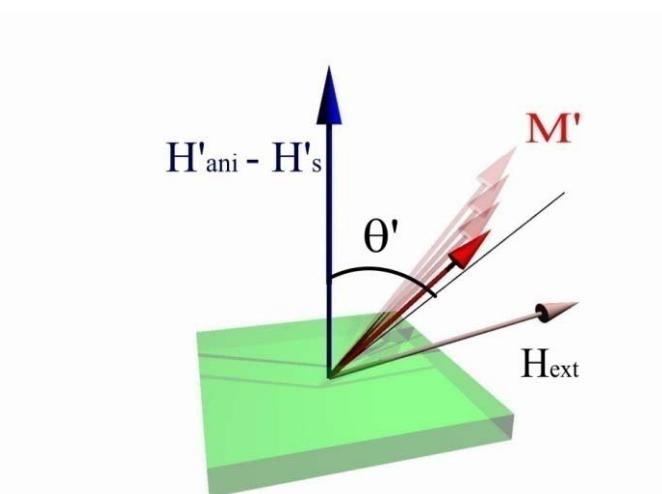


Optical pump-probe measurements of FMR

before pump pulse arrives

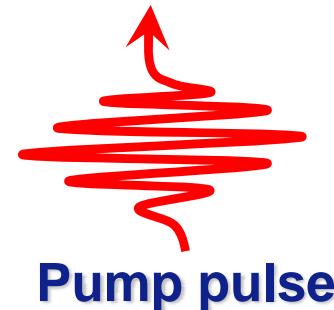


after pump has arrived

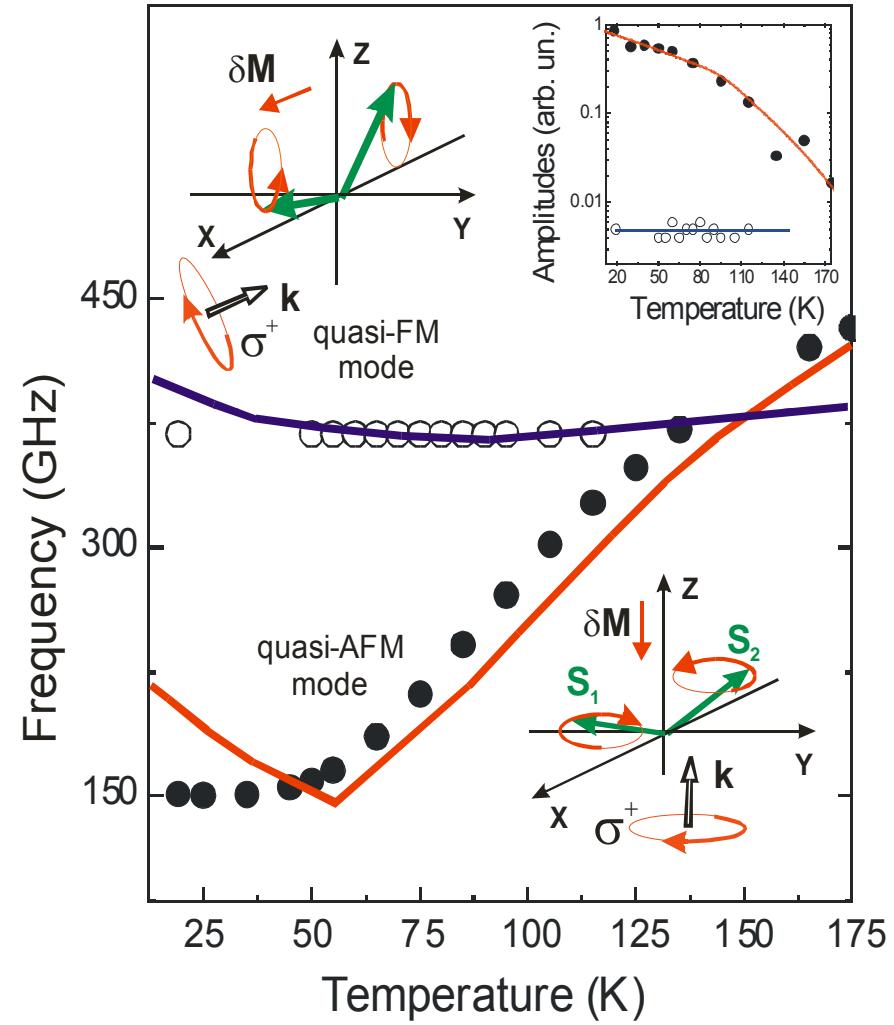
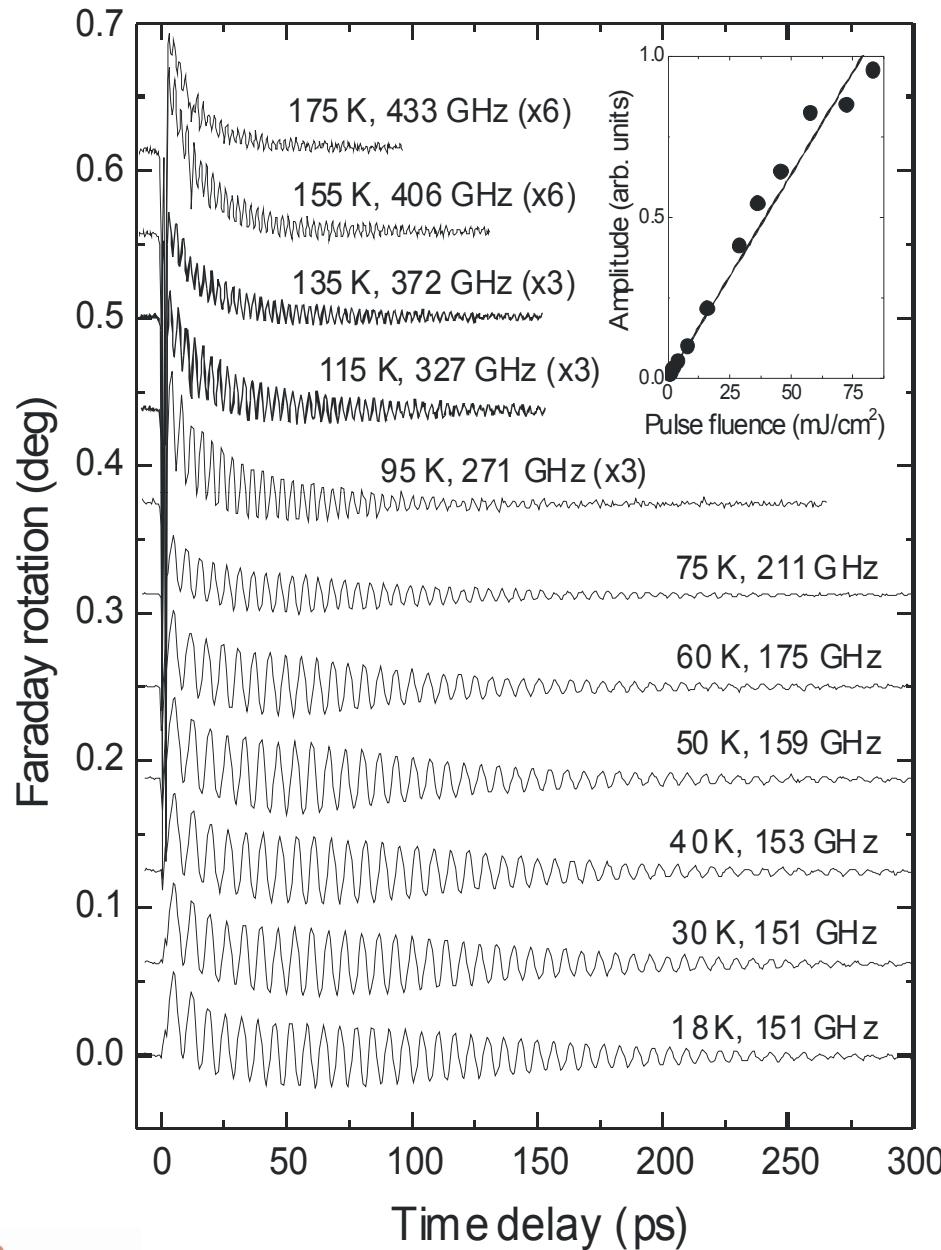


external field

$$\vec{H}^{eff} = \vec{H}_{ext} + \underbrace{\vec{H}_{ani} + \vec{H}_s}_{\text{affected by the laser}}$$



All-optical magnetic resonance in antiferromagnets



Outline of the lecture

- Light as a probe
 - linear magneto-optics
 - nonlinear (magneto-)optics
- Example: all-optical FMR
- Light as an excitation
 - classification of effects
 - basics of opto-magnetism
 - coherent control
 - local control of spins
- *can this become too-ultrafast?*

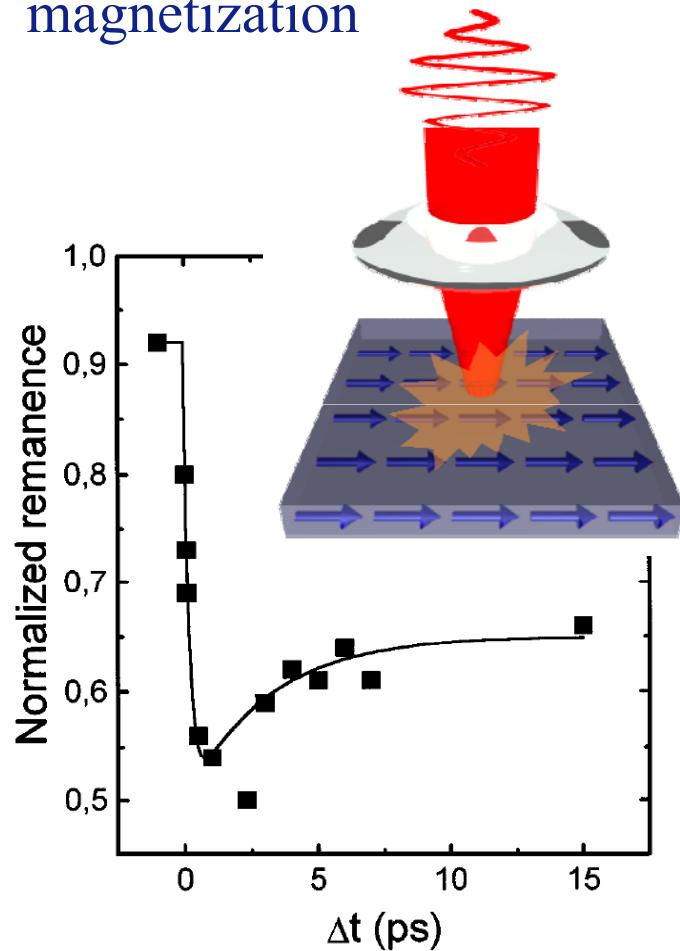
Effects of the laser pulse: classification

I. Thermal effects:

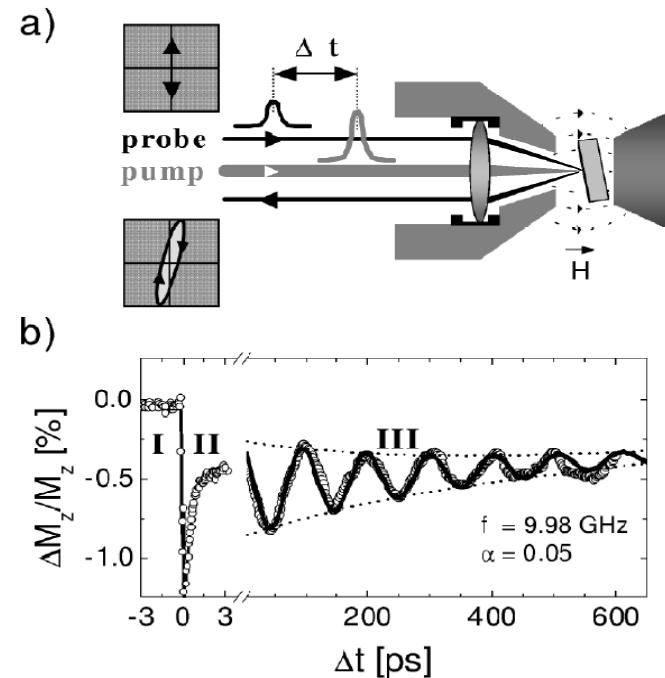
change of M is a result of change of T

Thermal laser-induced effects

laser-induced collapse of magnetization



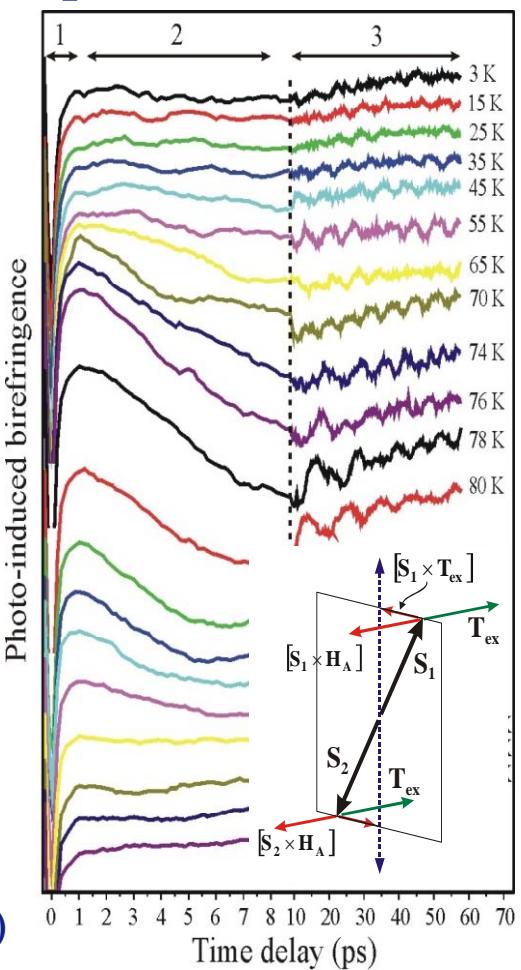
excitation and study
of spin waves



Ju et al., PRL **82**, 3705 (1999)
van Kampen et al, PRL **88**, 227201 (2002)

Beaurepaire et al, PRL **76**, 4250 (1996)

ultrafast
phase transitions



Kimel et al., Nature **429**, 850 (2004)
Ju et al, PRL **93**, 197403 (2004)
Thiele et al, APL **85**, 2857 (2004)

Effects of the laser pulse: classification

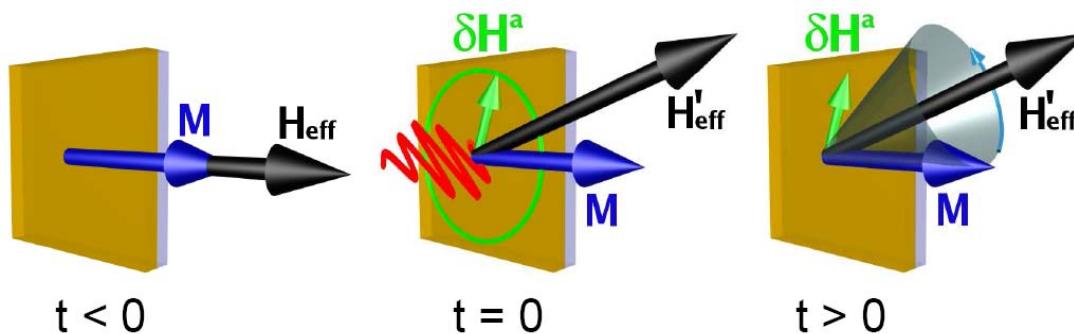
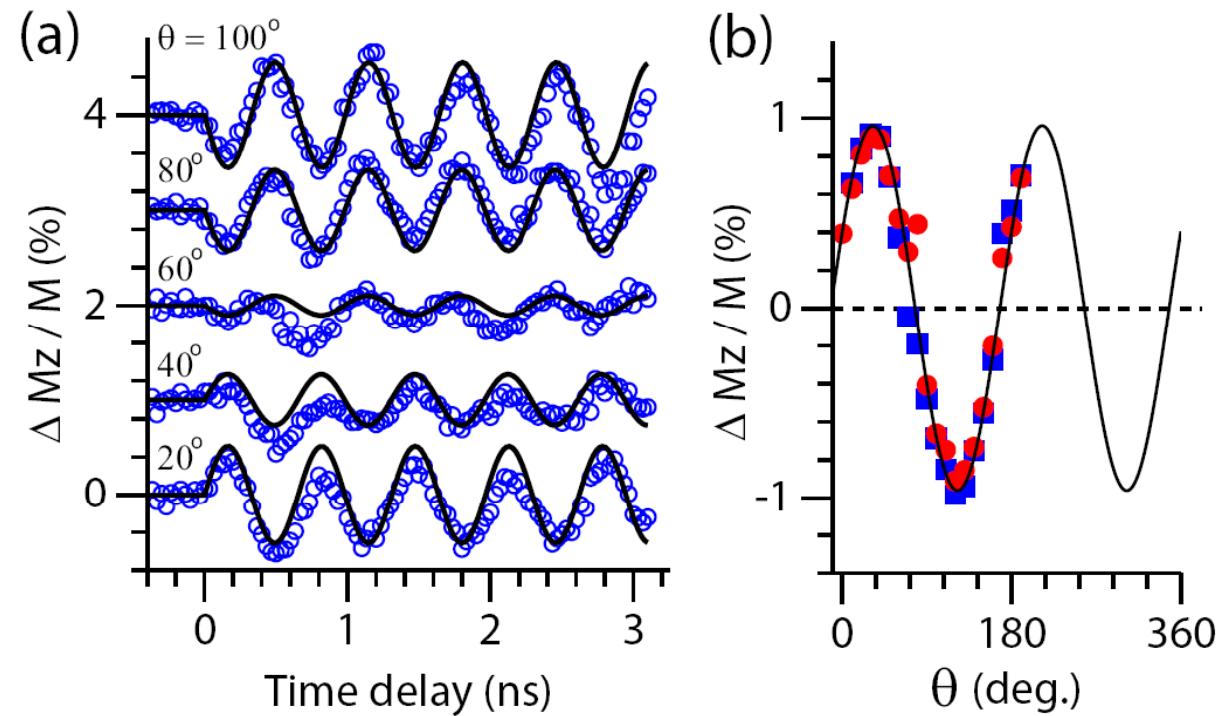
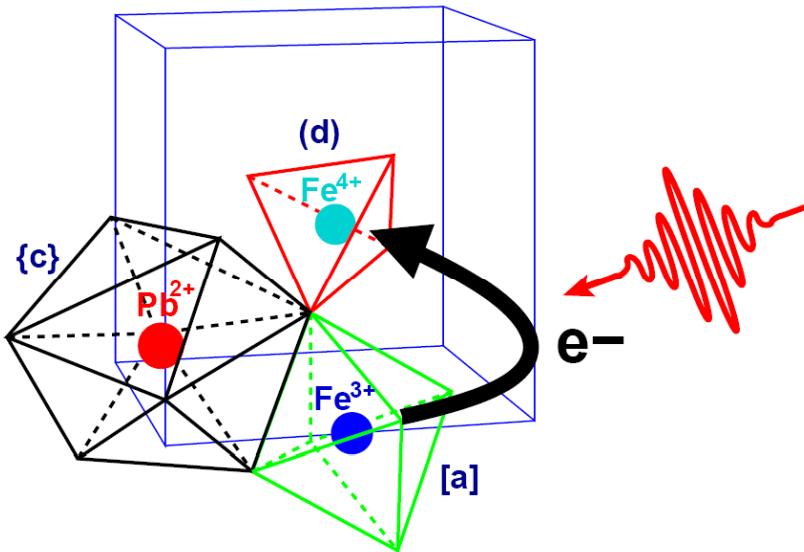
I. Thermal effects:

change of M is a result of change of T

II. Nonthermal photo-magnetic effects:

based on photon absorption

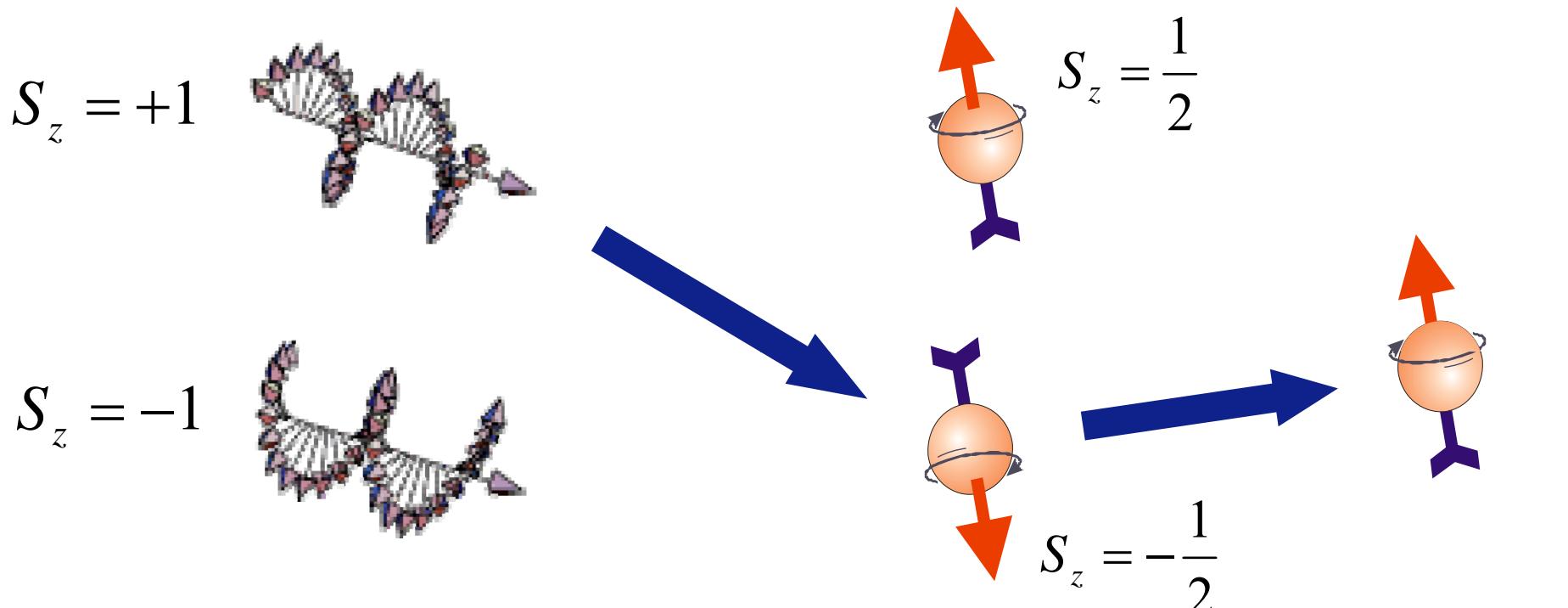
Photo-magnetic effects: modification of anisotropy



**polarization-dependent
effect => nonthermal!**

Hansteen *et al.*, PRL **95**, 047402 (2005);
Phys. Rev. B **73**, 014421 (2006).

Circular polarization, photon spin, and absorption



$$1 \text{ photon / site} = 20000 \text{ K } \Delta T$$

$\sim 0.01 \text{ phot/site max}$
 $\leq 0.01 \text{ efficiency}$

} effect $\leq 10^{-4}$

very fast and easy?



Effects of the laser pulse: classification

I. Thermal effects:

change of M is a result of change of T

II. Nonthermal photo-magnetic effects:

based on photon absorption

III. Nonthermal opto-magnetic effects:

do not require absorption

Extended introduction in laser-induced dynamics

REVIEWS OF MODERN PHYSICS, VOLUME 82, JULY–SEPTEMBER 2010

Ultrafast optical manipulation of magnetic order

Andrei Kirilyuk,^{*} Alexey V. Kimel, and Theo Rasing

Institute for Molecules and Materials, Radboud University Nijmegen, Heyendaalseweg 135, 6525 AJ Nijmegen, The Netherlands

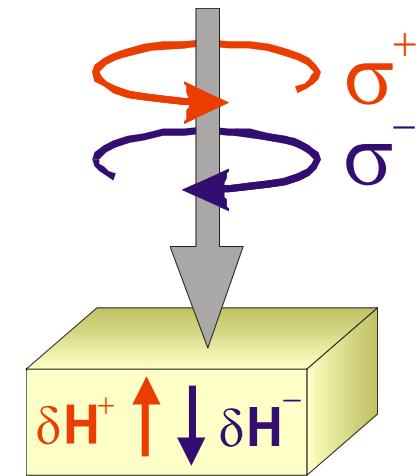
**everything you ever wanted to know about
laser-induced magnetization dynamics...**

Thermodynamics of magneto-optics

$$\Phi = \epsilon \epsilon_0 E(\omega) E^*(\omega)$$

$$H(0) = -\frac{1}{\mu_0} \frac{\partial \Phi}{\partial M(0)} = -\frac{\epsilon_0}{\mu_0} E(\omega) E^*(\omega) \frac{\partial \epsilon}{\partial M}$$

$$\hat{\epsilon} = \begin{pmatrix} \epsilon_{xx} & -i\alpha M & 0 \\ +i\alpha M & \epsilon_{yy} & 0 \\ 0 & 0 & \epsilon_{zz} + O(M^2) \end{pmatrix}$$



Inverse Faraday effect

$$\vec{H}(0) = \frac{\epsilon_0}{\mu_0} \alpha [\vec{E}(\omega) \times \vec{E}^*(\omega)]$$

Pitaevskii, Sov. Phys. JETP **12**, 1008 (1961).
van der Ziel Phys. Rev. Lett. **15**, 190 (1965).

Inverse Faraday effect

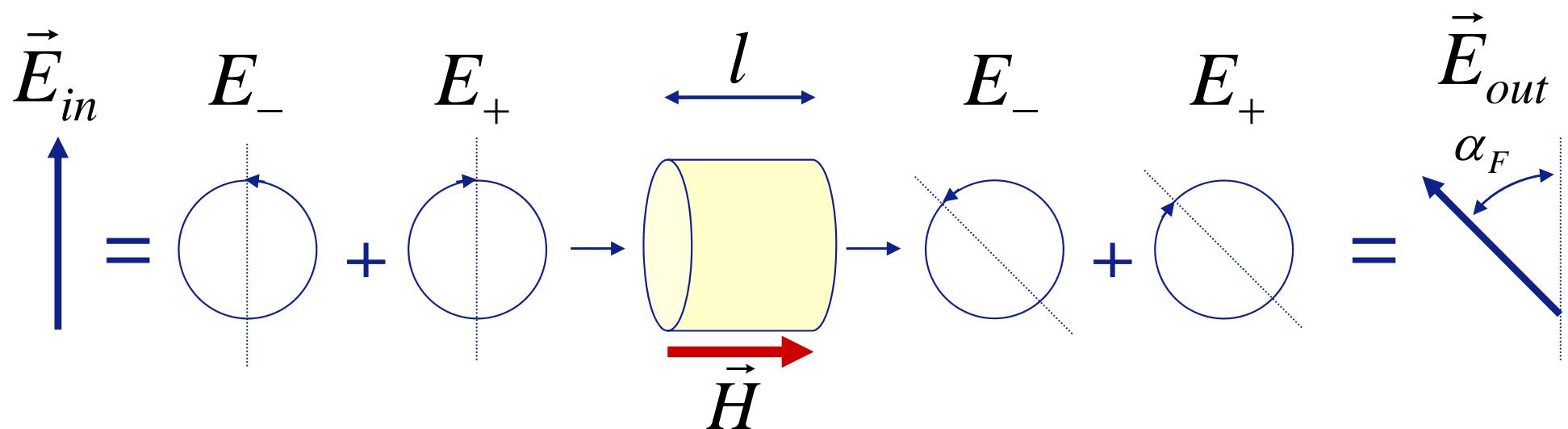
$$\text{Faraday rotation: } \theta_F = \frac{2\pi l}{\lambda} \frac{\alpha M}{\epsilon_0}$$

$$\vec{H}(0) = \frac{\epsilon_0}{\mu_0} \alpha [\vec{E}(\omega) \times \vec{E}^*(\omega)]$$

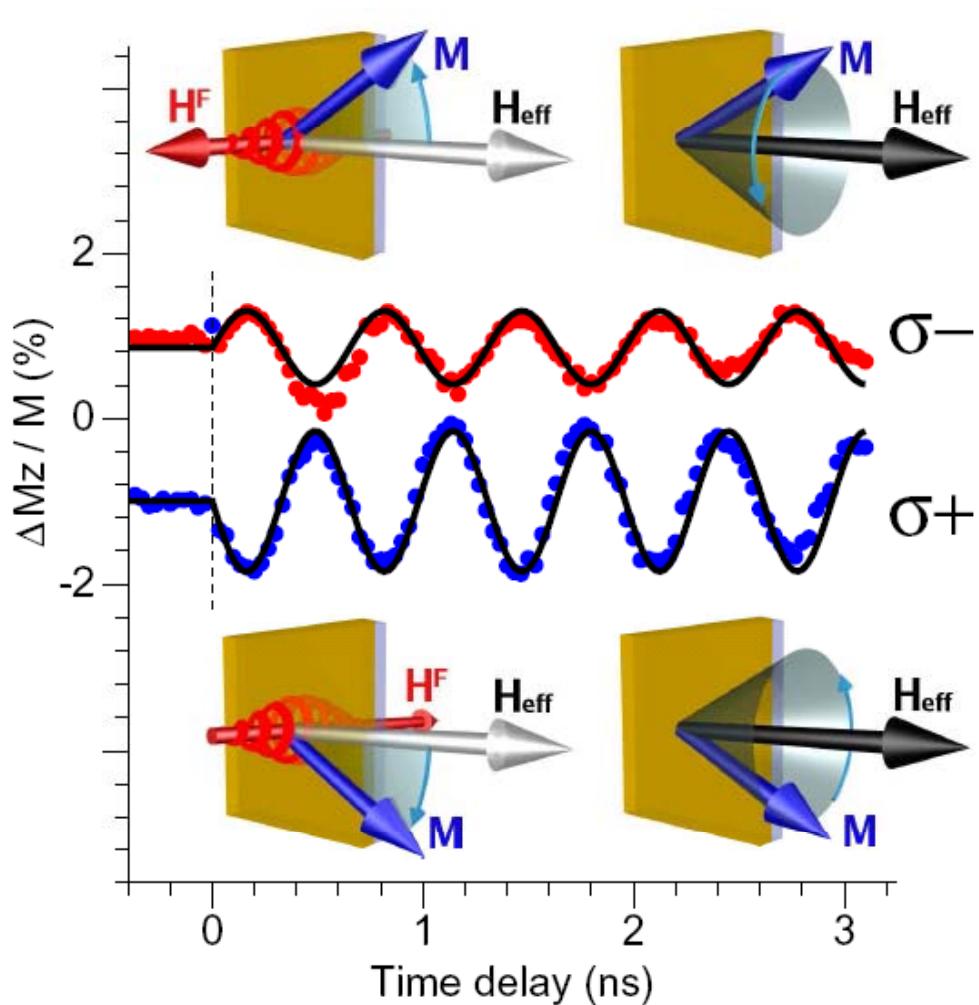


no absorption required!

no angular momentum transfer!



Effect for opposite pulse helicities



→ it works!!!

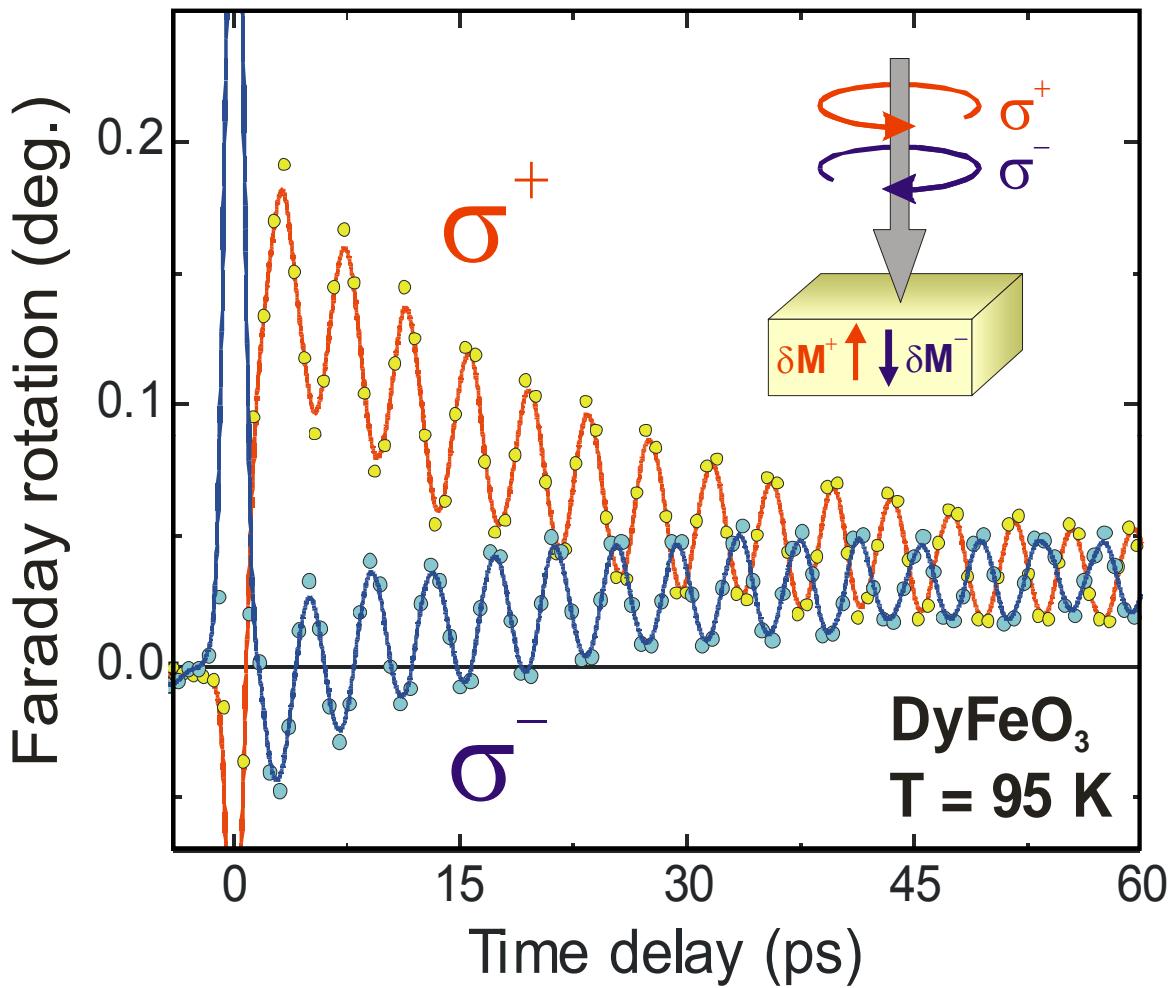
equivalent to a 100 fs
magnetic field pulse of
some 0.5–1 Tesla!

$$\vec{H}(0) = \frac{\epsilon_0}{\mu_0} \alpha [\vec{E}(\omega) \times \vec{E}^*(\omega)]$$

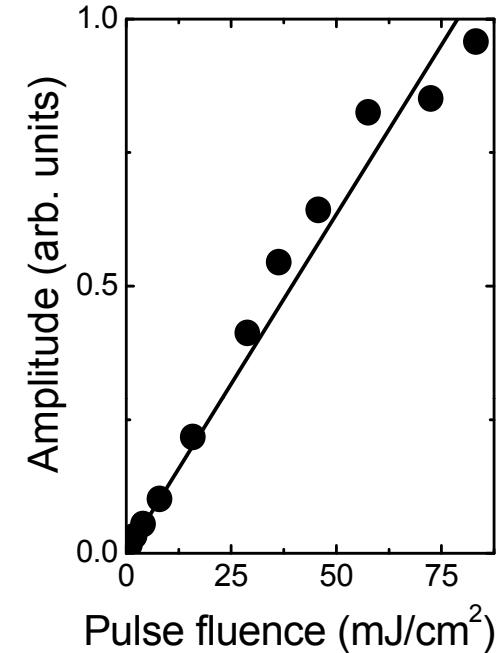
$$H_{IFE} \sim 0.1\text{--}100 \text{ Tesla}$$

Hansteen *et al.*, PRL 95, 047402 (2005);
Phys. Rev. B 73, 014421 (2006).

Works everywhere! (almost)



Kimel *et al.*, Nature **435**, 655 (2005)

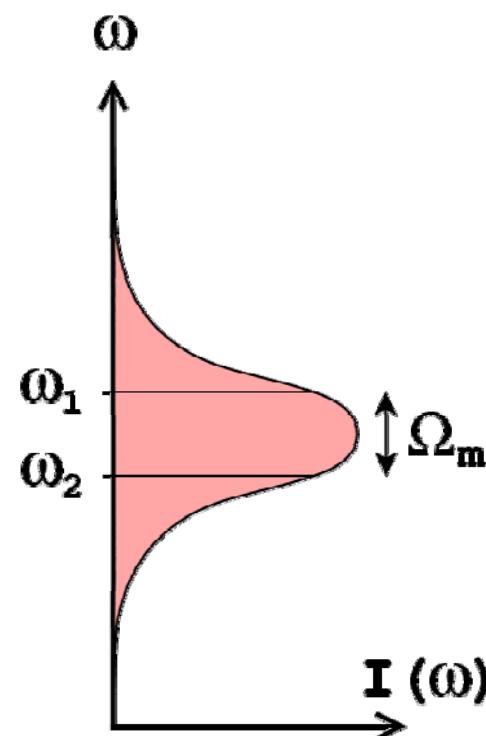
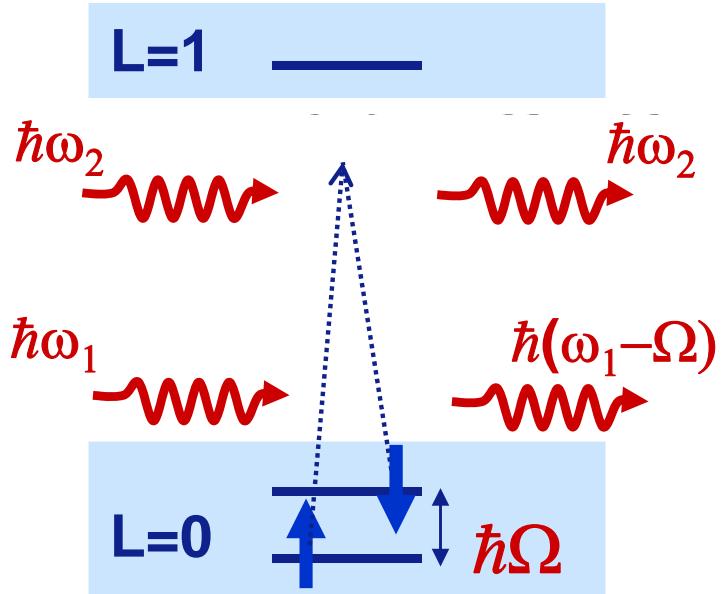


$$\vec{H}(0) = \frac{\epsilon_0}{\mu_0} \alpha [\vec{E}(\omega) \times \vec{E}^*(\omega)]$$

Microscopic mechanism of the inverse Faraday effect

- Stimulated Raman scattering on magnons (2-photon process)

[Shen et al, Phys. Rev. (1966)]

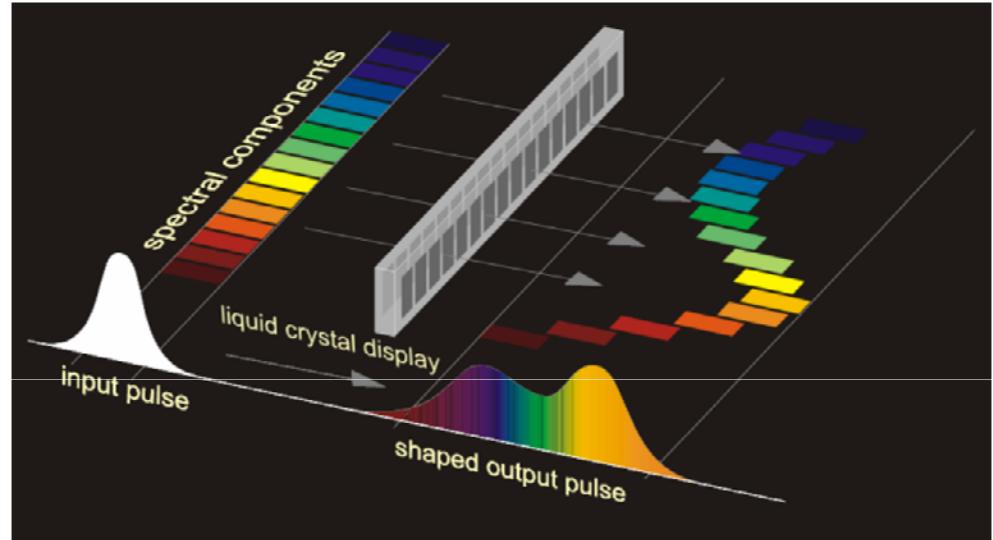
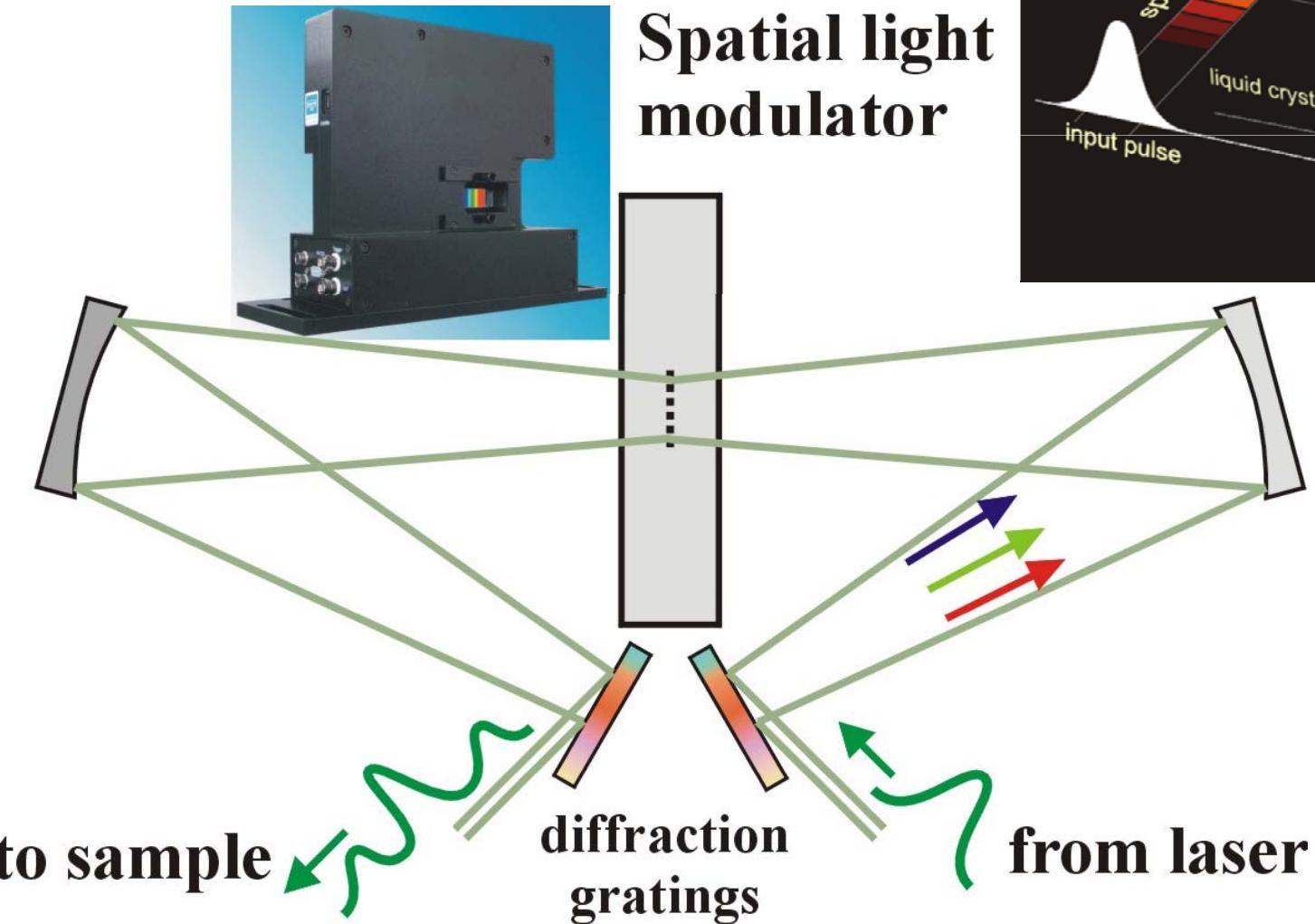


- Number of photons is conserved
- Process can be fast
 $\tau \sim 1 / \omega \sim 1 \text{ fs}$

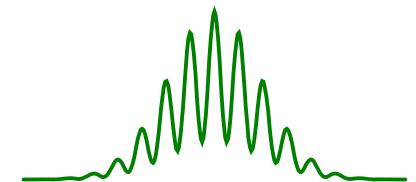
light helicity (= angular momentum) is also conserved!

Manipulating pulse frequencies

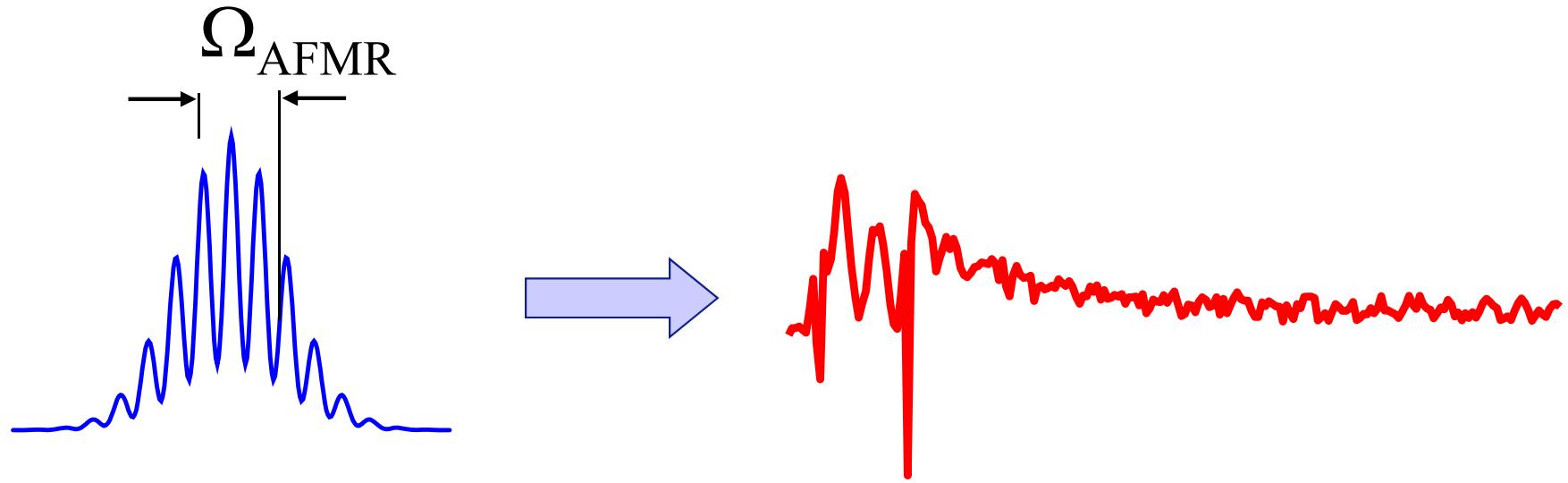
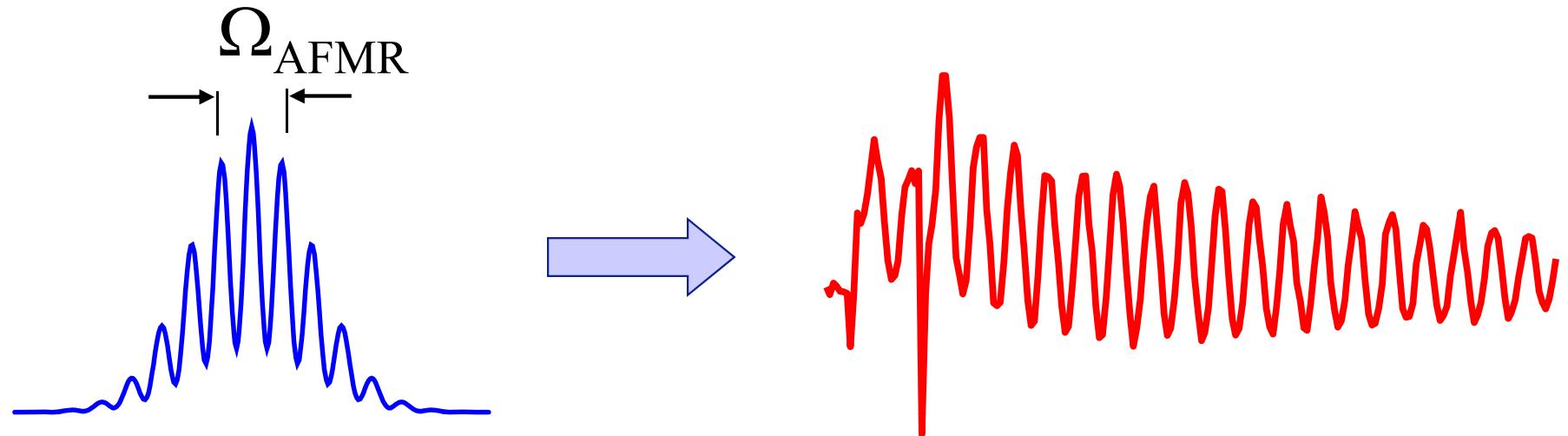
picture courtesy Th.Baumert



Amplitudes and phases
of 320 components



Opto-magnetic effect with “shaped” pulse

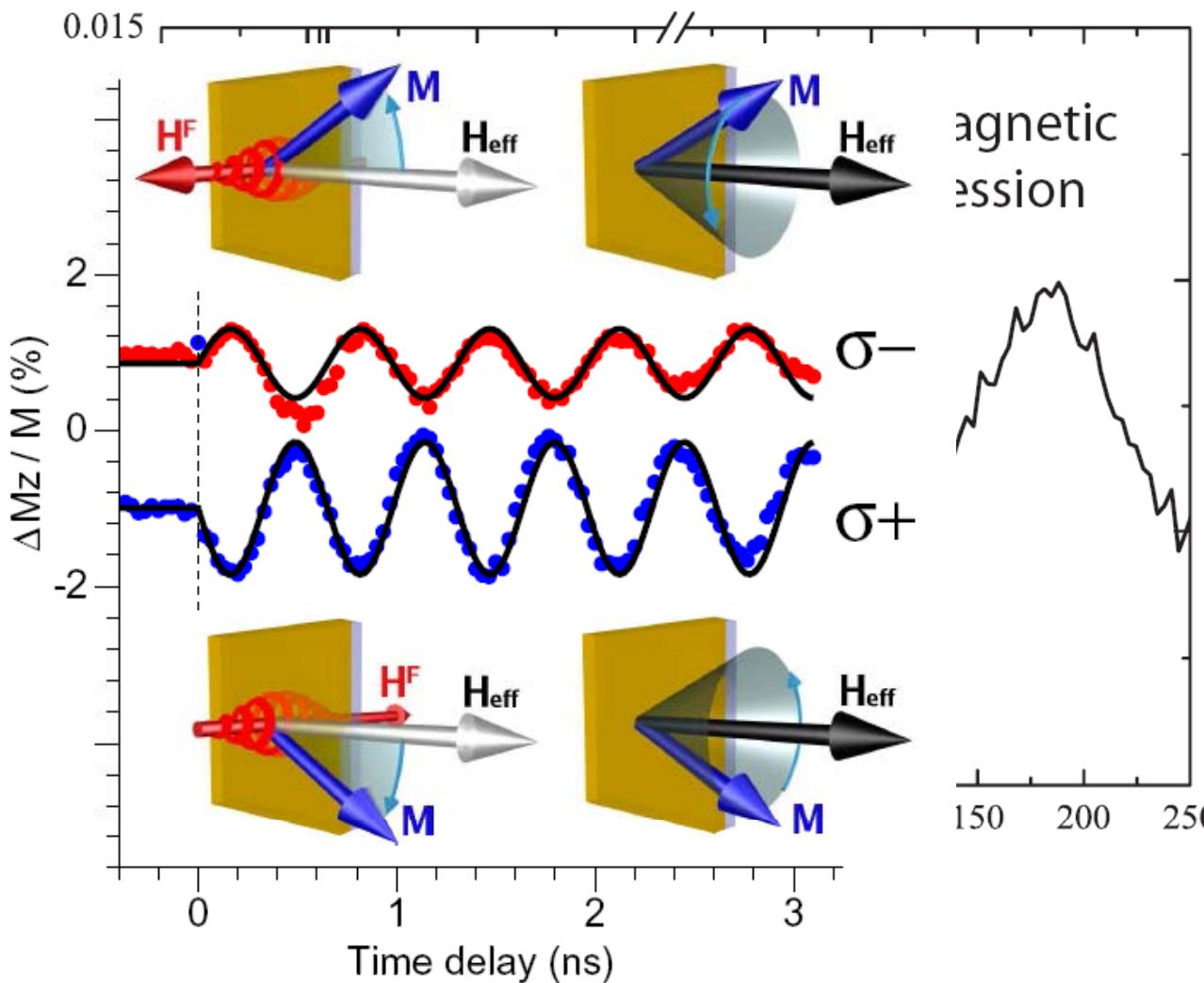


Outline of the lecture

- Light as a probe
 - linear magneto-optics
 - nonlinear (magneto-)optics
- Example: all-optical FMR
- Light as an excitation
 - classification of effects
 - basics of opto-magnetism
 - coherent control
 - local control of spins
- *can this become too-ultrafast?*

Higher frequency component?

A. Reid et al.,
Phys. Rev. Lett. **105**, 107402 (2010).

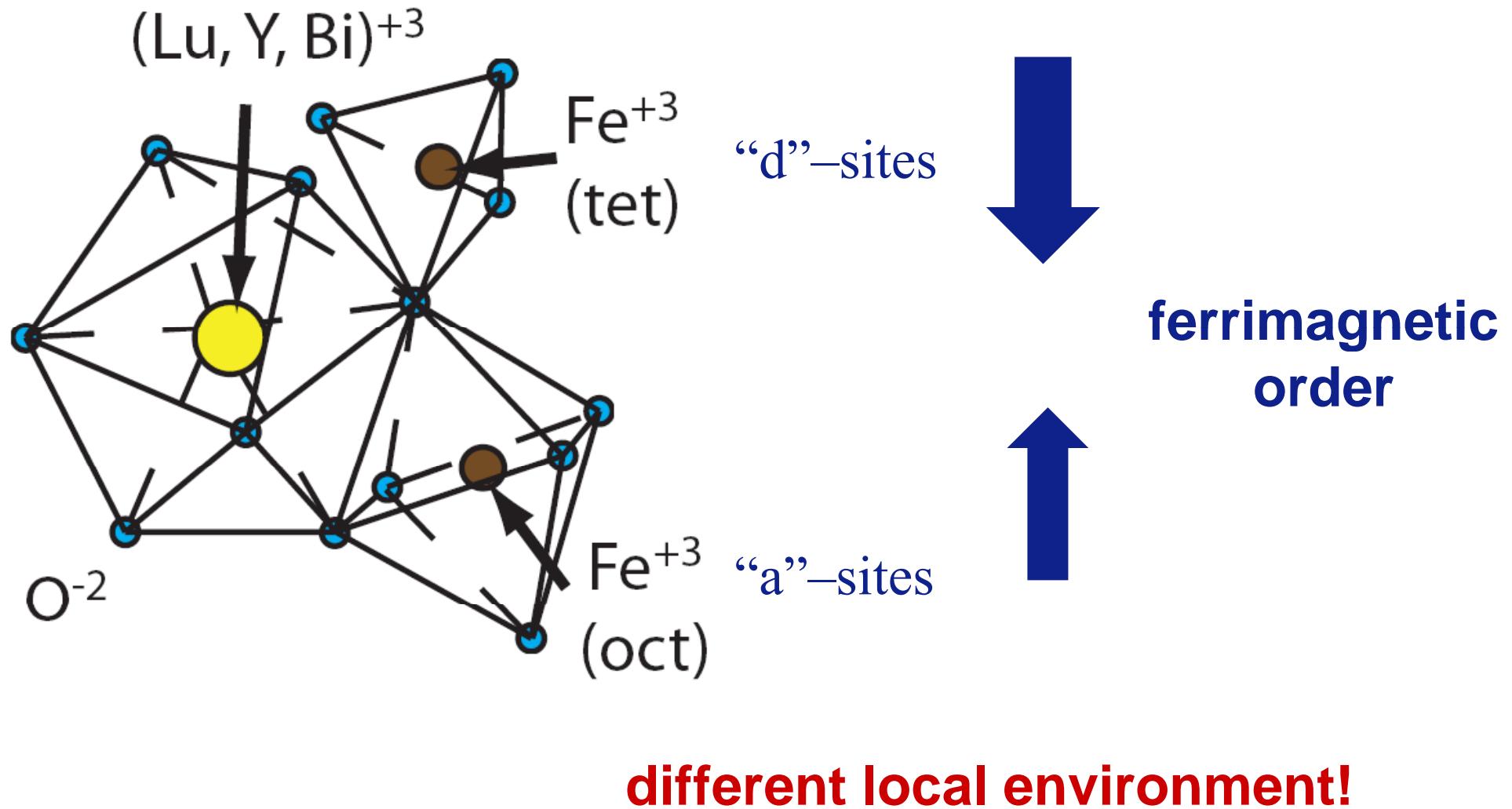


**High frequency:
650 GHz**
**Phase change
with pump helicity.**

Kaplan–Kittel exchange resonance

J. Kaplan and C. Kittel, *J. Chem Phys.* **21** 760-761 (1953).

Garnet structure $[\text{Lu}_{1.69}\text{Y}_{0.65}\text{Bi}_{0.66}](\text{Fe}_{3.85}\text{Ga}_{1.15})\text{O}_{12}$

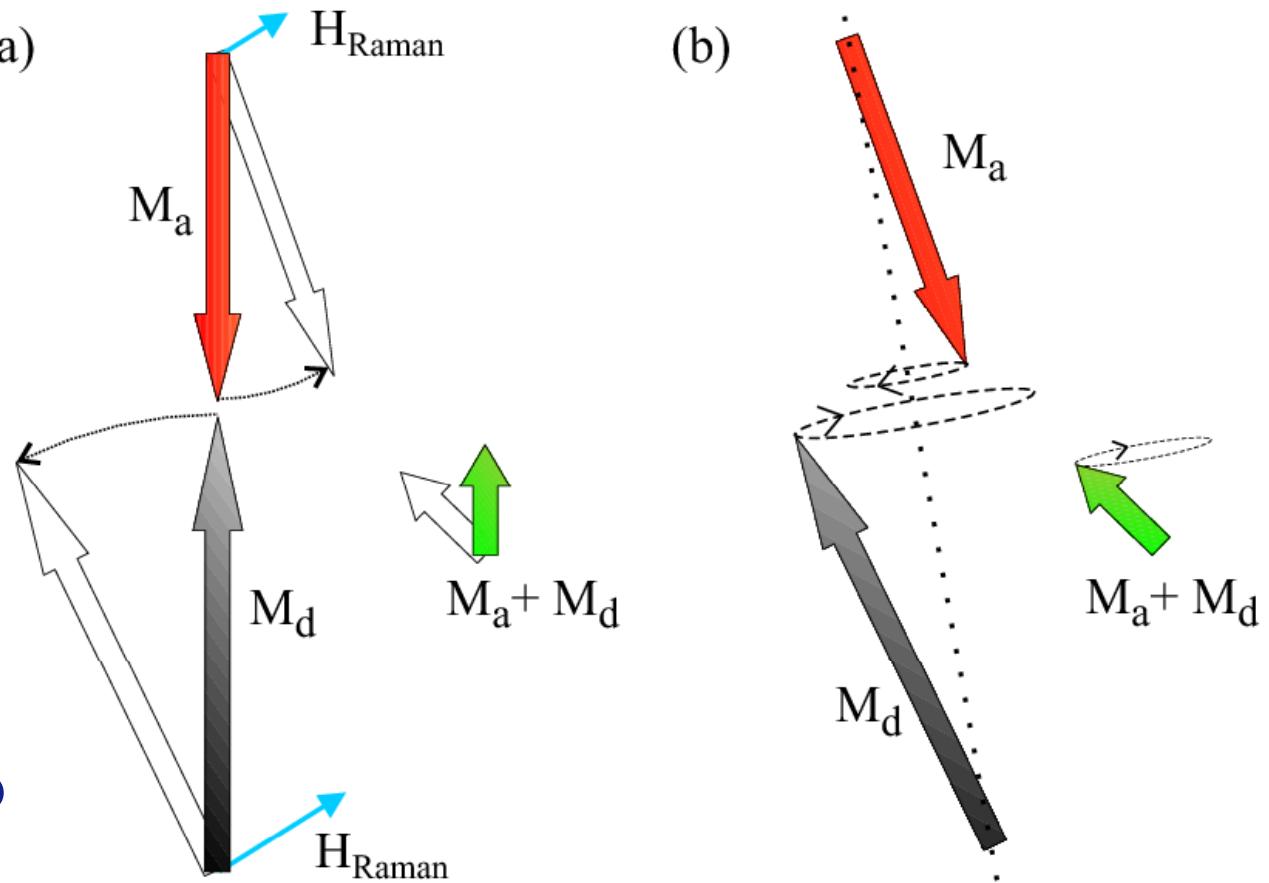


Exchange Resonance

$$\frac{d\mathbf{M}_a}{dt} = \gamma_a \mathbf{M}_a \times \mathbf{H}_a,$$

$$\frac{d\mathbf{M}_d}{dt} = \gamma_d \mathbf{M}_d \times \mathbf{H}_d,$$

Shouldn't be able to see it.



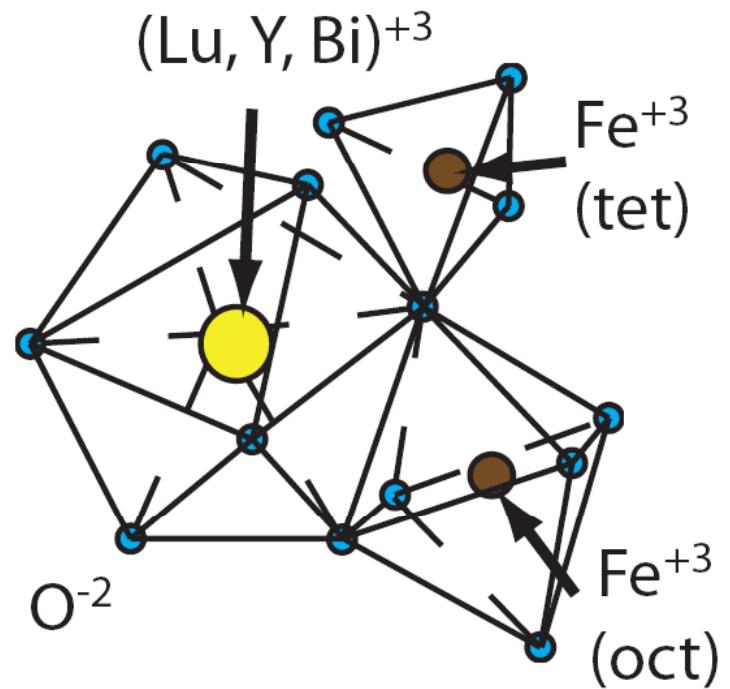
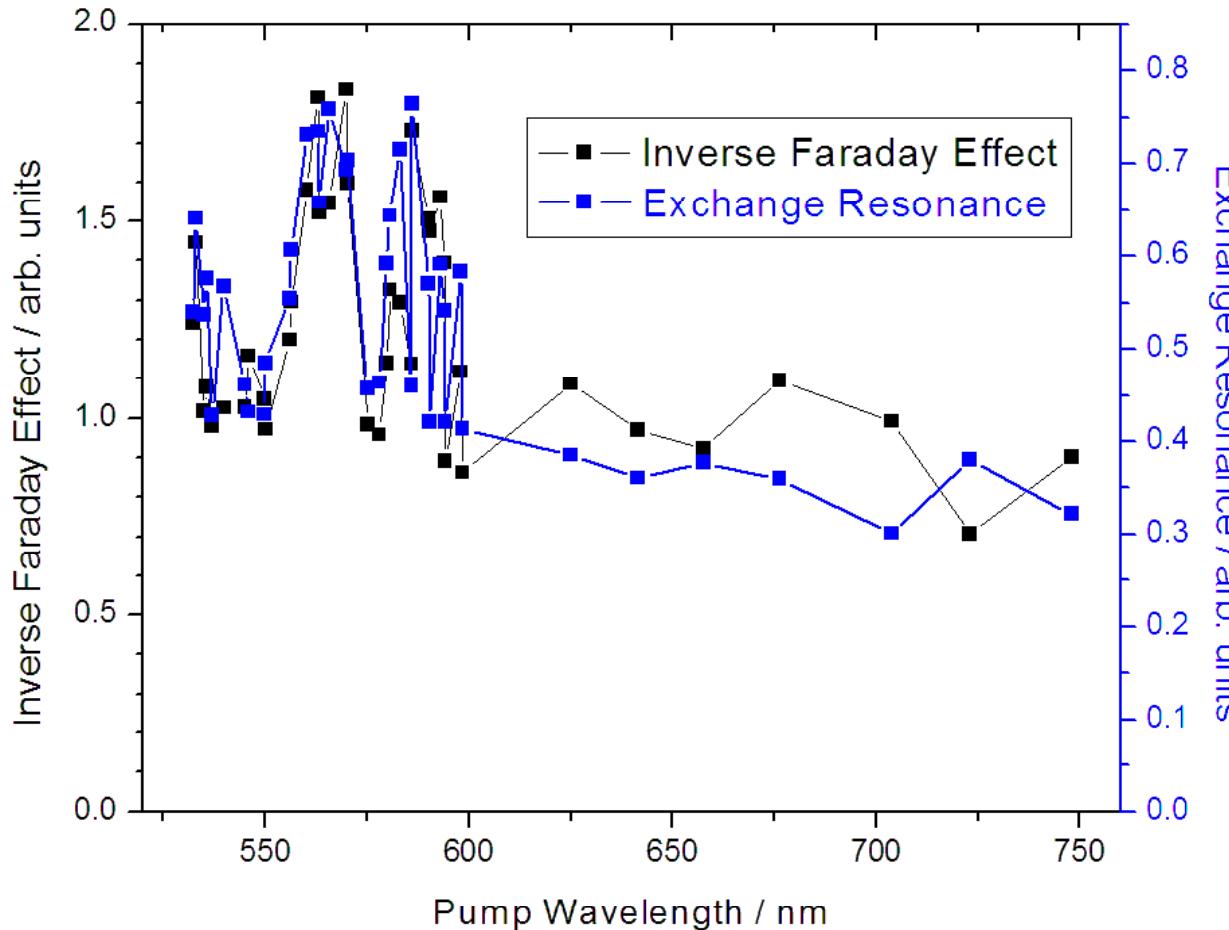
$$\omega_{\text{ex}} = |\gamma_a \lambda_{\text{ad}} \mathbf{M}_a - \gamma_d \lambda_{\text{ad}} \mathbf{M}_d|,$$

Neither OM nor MO necessarily correlate with \mathbf{M} .

Correlation with the IFE

A. Reid et al.,
Phys. Rev. Lett. **105**, 107402 (2010).

The same spectral dependence



locally driven spin dynamics!

Outline of the lecture

- Light as a probe ✓
 - linear magneto-optics ✓
 - nonlinear (magneto-)optics ✓
- Example: all-optical FMR ✓
- Light as an excitation ✓
 - classification of effects ✓
 - basics of opto-magnetism ✓
 - coherent control ✓
 - local control of spins ✓
- *can this become too-ultrafast?*

Transient magneto-optical response

Transient complex (Kerr or Faraday) rotation

$$\tilde{\theta}(t) = G(t) + F(t)M(t)$$

Pump-induced change $\Delta\tilde{\theta}_T(t) = \Delta G(t) + M_0\Delta F(t) + F_0\Delta M(t)$

If, by some chance $\begin{cases} F(t) \equiv F_0 \\ G(t) \equiv G_0 \end{cases}$, then $\tilde{\theta}(t) = G_0 + F_0M(t)$

and

$$\frac{\Delta\theta(t)}{\theta_0} = \frac{\Delta\varepsilon(t)}{\varepsilon_0} = \frac{\Delta M(t)}{M_0}$$

Ultrafast Magneto-Optics in Nickel: Magnetism or Optics?

B. Koopmans,* M. van Kampen, J. T. Kohlhepp, and W. J. M. de Jonge

*Eindhoven University of Technology, Department of Applied Physics, COBRA Research Institute,
P.O. Box 513, 5600 MB, Eindhoven, The Netherlands*

(Received 22 February 2000)

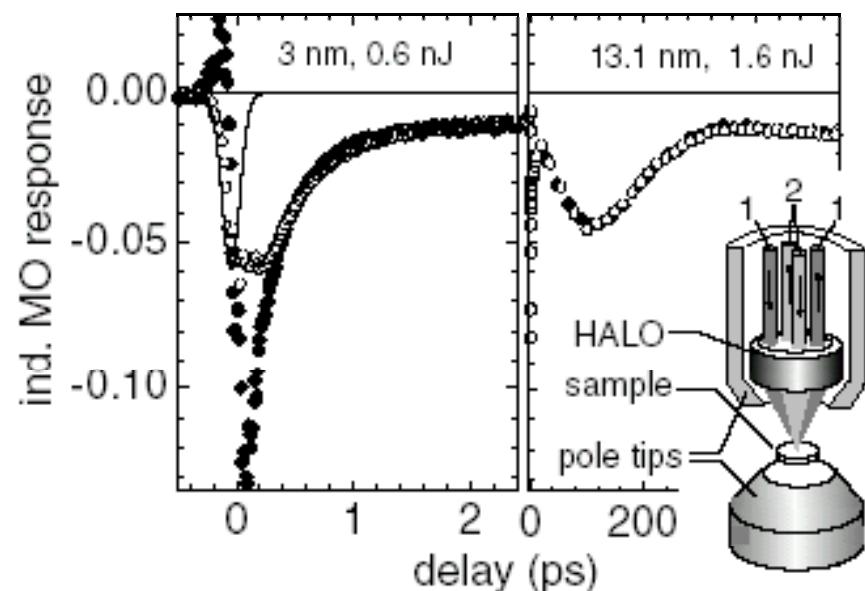


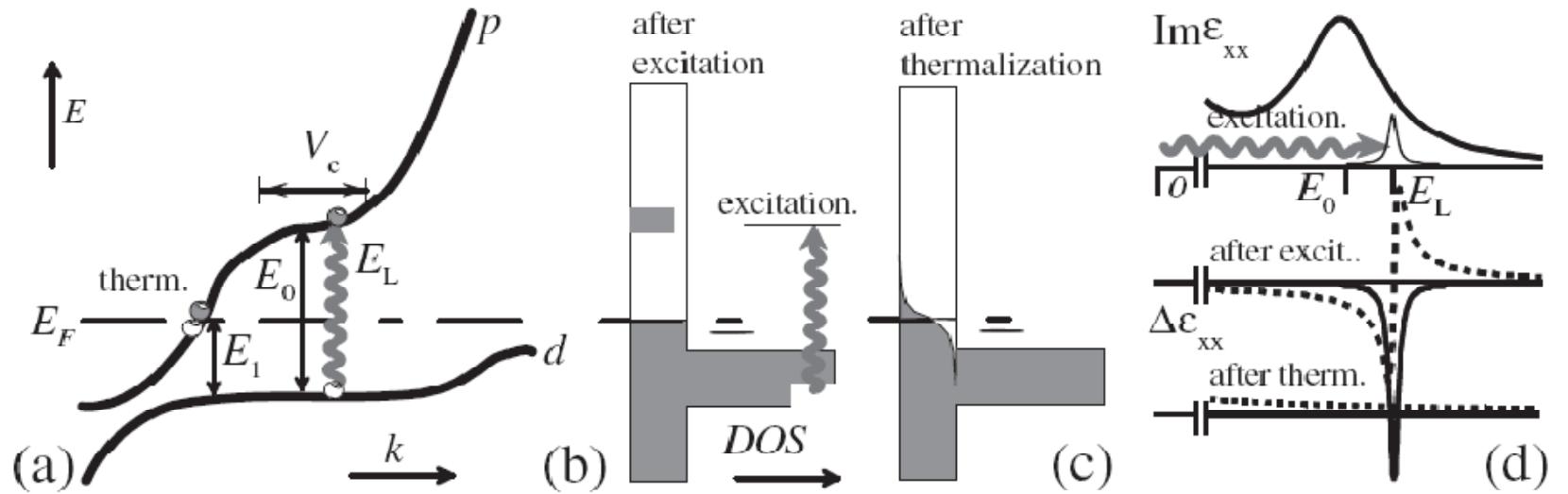
FIG. 1. Comparison of the induced ellipticity ($\Delta\psi''/\psi_0''$, open circles) and rotation ($\Delta\psi'/\psi_0'$, filled diamonds) as a function of pump-probe delay time, for a (111) oriented film at the thicknesses and pulse energies indicated. The thick line represents the pump-probe cross correlation trace. The inset depicts the experimental configuration with pump ("1") and probe ("2") beams.

$$\frac{\Delta\theta(t)}{\theta_0} = \frac{\Delta\epsilon(t)}{\epsilon_0} = \frac{\Delta M(t)}{M_0}$$

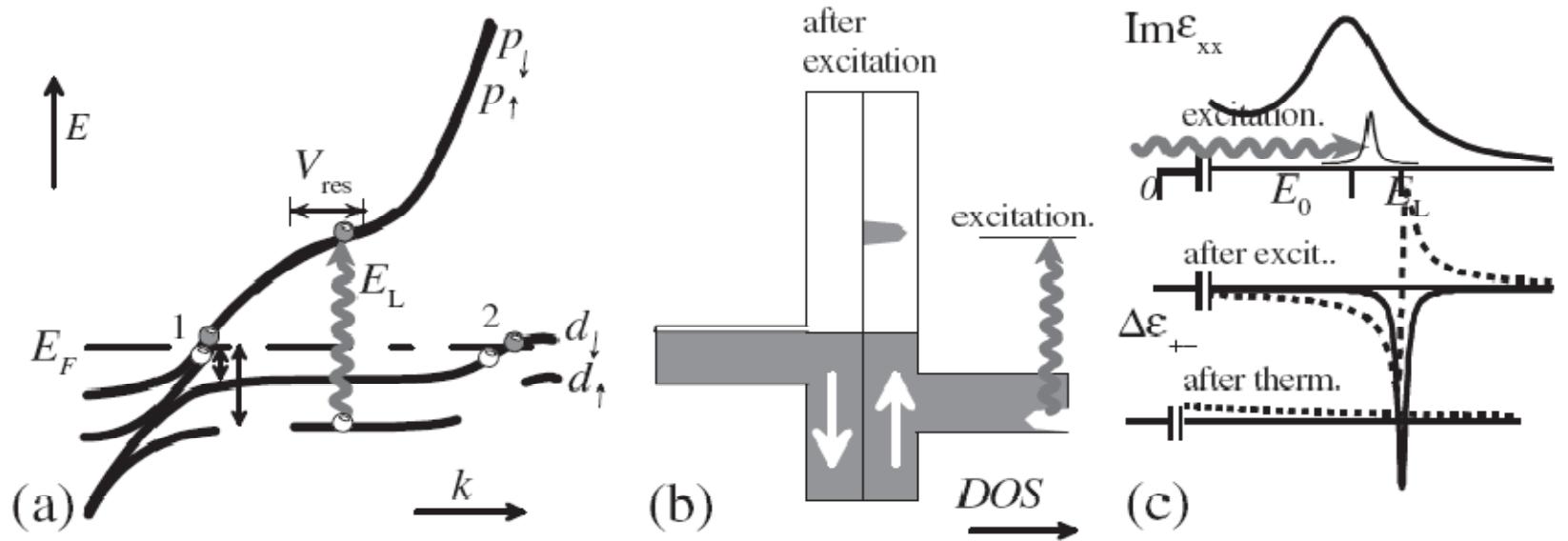
system out of equilibrium

Optical effects?

nonmagnetic



magnetic



Messages to take home

- not everything you measure is magnetization
- opto-magnetic effects lead to real change of M during the pulse
- it is a challenge to show whether there is any other nonthermal mechanism to do this!